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Polarization-dependent ponderomotive gradient force in a standing wave

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The ponderomotive force is derived for a relativistic charged particle entering an electromagnetic standing wave with a general three-dimensional field distribution and a nonrelativistic intensity, using a perturbation expansion method. It is shown that the well-known ponderomotive gradient force expression does not hold for this situation. The modified expression is still of simple gradient form but contains additional polarization-dependent terms. These terms arise because the relativistic translational velocity induces a quiver motion in the direction of the magnetic force, which is the direction of large field gradients. Consistent perturbation expansion of the equation of motion leads to an effective doubling of this magnetic contribution. The derived ponderomotive force generalizes the polarization-dependent electron motion in a standing wave obtained earlier [A. E. Kaplan and A. L. Pokrovsky, Phys. Rev. Lett. 95, 053601 (2005)]. Comparison with simulations in the case of a realistic, nonidealized, three-dimensional field configuration confirms the general validity of the analytical results.

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I. INTRODUCTION

The ponderomotive force is a time-averaged force experienced by a charged particle in an oscillating electromagnetic (EM) field that is spatially inhomogeneous. In the standard perturbative approach [1,2], it is shown that a charged particle in an oscillating EM field attains an oscillatory quiver momentum superimposed on a slowly varying guiding-center momentum \( \mathbf{p} \). The latter is subject to the classical ponderomotive force \( F_p \):

\[
\frac{d\mathbf{p}}{dt} = F_p = -\frac{e^2}{2\varepsilon_0mc\omega^2} \nabla I(x),
\]

where \( m \) is the mass of the particle, \( e \) its charge, \( \varepsilon_0 \) the permittivity, \( \omega \) the frequency of the EM field, and \( I(x) \) the position-dependent field intensity. The classical ponderomotive force is of gradient form and always directed toward regions of low field intensity. The ponderomotive force is observed and exploited in a wide range of contexts. In laser-plasma physics, the ponderomotive force drives the formation of laser wakefields that are used for next generation electron accelerators [3,4]. Ion beams are produced by intense laser irradiation of thin foils, in which the ponderomotive force plays an essential role [5,6]. Schemes have been proposed for ponderomotive laser-vacuum acceleration of electrons [7,8]. In Paul traps, ions are confined by a ponderomotive potential [9]. In electron beam diagnostics, the length of electron bunches is measured by sequentially scattering different sections of the bunch using the ponderomotive force of a laser pulse [10,11].

The field gradients that can be obtained in a single laser pulse are set by the laser pulse duration longitudinally and the focal spot transversely. For many applications of the ponderomotive force this means that, to obtain a sufficiently strong force, field intensities are required that are high enough to cause a relativistic quiver motion (which happens if the normalized amplitude of the vector potential, \( a = eA/mc = e\sqrt{2I/(\varepsilon_0c)/(mc\omega)} \gtrsim 1 \); that is, \( I \gtrsim 2 \times 10^{18} \) W/cm\(^2\) for a wavelength of 800 nm). Relativistic field intensities necessitate more complicated descriptions of the average EM force [12–14] or, at least, restrict the domain of validity of Eq. (1) [15,16]. An intermediate situation occurs when an already relativistic particle enters an EM field with nonrelativistic intensity. A relativistic derivation [17] shows that this introduces an additional factor \( \sqrt{1 + \mathbf{p}^2/(mc)^2 + a^2/2} \) in the denominator of Eq. (1), resulting in an accurate description for practical situations [18].

An alternative to the application of a relativistic laser pulse is the use of a standing wave. In this configuration the nodes and antinodes are spaced on the scale of the wavelength, resulting in large field gradients. For example, a standing wave produced by two counterpropagating EM waves with wavelength \( \lambda = 800 \) nm and the very modest, nonrelativistic peak field intensity of \( 10^{15} \) W/cm\(^2\) already causes ponderomotive forces of the order of \( F_p/e \sim 1 \) GV/m.

For this reason, a number of applications of the ponderomotive force have been proposed that take advantage of the large field gradients in a standing EM wave. Hebeisen et al. [19] suggested a tabletop standing-wave version of the bunch length measurement setup mentioned previously. Following an earlier idea [20], Balcu proposed a novel x-ray-free electron laser based on the wiggling of electrons in the ponderomotive potential of a standing wave [21]. Faure et al. used a standing wave formed by colliding laser pulses to preaccelerate electrons ponderomotively in a laser-wakefield setup [22], demonstrating that the production of monoenergetic electron beams can be made stable and reproducible in that way [23–25]. Baum and Zewail proposed to create attosecond electron pulse trains by bunching of an electron beam due to a comoving ponderomotive beat potential between laser pulses of different frequencies [26].

In view of all these important technological applications, a thorough understanding of the ponderomotive force in a standing wave is essential. The scattering of charged particles by a standing EM wave was first described in a quantum-mechanical context by Kapitza and Dirac [27], and since

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then many papers have appeared on this subject [28–32]. Nevertheless, there are only a few classical electrodynamical studies on the standing-wave ponderomotive force [33–36].

Most publications on the ponderomotive force have concentrated on propagating EM waves, establishing the validity of Eq. (1) in that context [1,2,17]. Equation (1) is also applied to the standing-wave configuration, implicitly assuming that it remains valid in that case as well. In 2005, however, Kaplan and Pokrovsky [37,38] calculated the time-averaged equation of motion of an electron in a standing wave for a number of field polarizations, and their results showed that the ponderomotive force depends on the polarization. Most notably, the ponderomotive force can even change its direction toward high field regions for certain situations. Clearly, these results are in conflict with the polarization-independent Eq. (1), which is commonly used. Kaplan and Pokrovsky did not provide an alternative expression for the ponderomotive force, however.

We would first like to show, on the basis of simple arguments, that it can be understood that the ponderomotive force in a standing wave is polarization dependent. Consider Fig. 1, showing a particle with charge $e$ and initial velocity $v_0$ parallel to the $x$ axis, incident on a standing wave with electric field $E$ and magnetic field $B$. The wave is oriented with its nodal planes parallel to the $(x,y)$ plane, so that the spatial variation of the field is much more rapid in the $z$ direction than in the transverse direction. This is the typical system considered in this paper. When the particle enters the EM field, it will start to quiver in the polarization direction in response to the oscillating electric force $eE$. This electric quiver, combined with the Lorentz force equation, leads to the well-known average force, Eq. (1), independent of the polarization direction. However, the incident particle will also quiver in response to the magnetic force $\approx e(v_0 \times B)$. This magnetic quiver can be comparable in magnitude to the electric quiver for relativistic particles. Because the magnetic quiver is in the $z$ direction, the particle samples a large field gradient, leading to an additional contribution to the ponderomotive force that is comparable to the electric one.

And since the magnetic force and hence the amplitude of the magnetic quiver depend on the angle between $v_0$ and $B$, the magnetic contribution is dependent on the polarization direction.

In this paper, the ponderomotive force is derived for a relativistic particle entering a nonrelativistic standing wave with a general three-dimensional field distribution. It is shown that indeed Eq. (1), or the relativistic equivalent, does not hold for this situation. This may have important implications for experiments and proposals based on the standing-wave ponderomotive force. The main result of this paper, Eq. (21), shows that the modified ponderomotive force is still of simple gradient form but contains additional polarization-dependent terms. We thus generalize the results of Kaplan and Pokrovsky, which follow naturally from our ponderomotive force expression. This paper is structured as follows. In Sec. II, the polarization-dependent ponderomotive gradient force in a standing wave, Eq. (21), is derived using a perturbation expansion method. The origin of the additional polarization-dependent terms is discussed. Next, Eq. (21) is applied in Sec. III to calculate the averaged equation of motion of an electron in specific standing-wave geometries, reproducing the results obtained by Kaplan and Pokrovsky. In Sec. IV, we validate our ponderomotive force expression by testing it against numerical simulations of electron trajectories in a realistic, nonidealized field configuration.

II. THE POLARIZATION-DEPENDENT PONDEROMOTIVE FORCE

A. Assumptions

Consider first the idealized case of two plane EM waves of equal frequency $\omega$ that counterpropagate along the $z$ axis and add to form a standing wave. In the Coulomb gauge, the vector potential $A$ of this ideal standing wave is then purely harmonic in time $t$ and position $z$; that is, $(\partial^2/\partial t^2 + \omega^2)A = 0$ and $(\partial^2/\partial z^2 + k^2)A = 0$, where $k = \omega/c$, with $c$ the speed of light. Furthermore, the vector potential satisfies $A_{\parallel} = 0$ and $\nabla \times A_{\parallel} = 0$, where $i = x, y$.

In practical applications, however, an EM standing wave differs from this idealized situation in two ways. First, the standing wave has a finite transverse extent, leading to a small but nonzero transverse gradient $\nabla_{\perp} A_{\parallel}$ and a small longitudinal component $A_z$. These two quantities are related by the gauge condition $\nabla \cdot A = 0$, and scale analysis of the latter shows that, symbolically,

$$A_z \sim \nabla_{\perp} A_{\parallel} \sim \epsilon. \quad (2)$$

Here, $\epsilon \ll 1$ is a small parameter measuring the magnitude of the field inhomogeneity and is used as the expansion parameter in the derivation that follows. For example, in Gaussian laser beams focused to a waist of size $w_0$, this parameter is $\epsilon \sim (kw_0)^{-1}$.

Second, the standing wave has a finite lifetime, so that the vector potential is only quasimonochromatic:

$$\frac{\partial^2 A}{\partial (ot)^2} + A = O(\delta A), \quad \frac{\partial^2 A}{\partial (kz)^2} + A = O(\delta A), \quad (3)$$
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where \( \delta \ll 1 \) is another small parameter measuring the monochromaticity. For a standing wave produced by counterpropagating laser pulses of temporal length \( \sigma \), for example, this parameter is \( \delta \sim (\omega \sigma)^{-1} \).

In addition, in this paper the EM field is assumed to be of nonrelativistic intensity, which means that

\[
\frac{eA}{mc} \equiv a \ll 1. \tag{4}
\]

For reasons of clarity, for the moment it is assumed that \( a \sim \epsilon \). However, the derivation below can be generalized straightforwardly to other field strengths such as \( a \sim \epsilon^2 \) and \( a \sim \epsilon^{1/2} \), leading to the same result. Appendix B gives a short description of the generalized derivation.

B. Perturbation expansions

Before considering the dynamics of a charged particle in a standing wave in detail, let us first determine what time scales are involved. First, there is the time scale of the quiver motion, which is the optical time scale \( \omega^{-1} \). Second, there is the time scale on which the motion of the guiding-center changes. Substituting an ideal standing wave \( A = A_0 \epsilon \cos kz \sin \omega t \) in Eq. (1) and integrating yields oscillatory motion in the \( z \) direction, with a typical frequency \( \Omega = \epsilon A_0 \omega / (\sqrt{2} mc) \). Thus the guiding-center motion in the \( z \) direction changes on a second, longer time scale \( \omega^{-1} \sim (\omega \sigma)^{-1} \). Finally, in a realistic standing wave, the nonzero transverse field gradient causes transverse ponderomotive forces, which, in view of Eqs. (1) and (2), are weaker than the longitudinal ponderomotive forces by a factor \( \epsilon \). Therefore the transverse guiding-center motion changes on a third, still longer time scale \( \epsilon^2 \omega^{-1} \).

Having established the three time scales of the problem, consider next the equations of motion of a charged particle in the standing wave [39]:

\[
\frac{d}{d(\omega t)} \left( \frac{p}{mc} + \frac{eA}{mc} \right) = \frac{1}{\gamma} \left( \frac{\lambda \nabla \cdot eA}{2mc} \right) - \frac{p}{mc}, \tag{5}
\]

\[
\frac{d(kx)}{d(\omega t)} = \frac{1}{\gamma mc}, \tag{6}
\]

in which \( \gamma = \sqrt{1 + p^2/(mc)^2} \) is the Lorentz factor, and the dyadic notation \( \nabla A \) has been used [40]. Equations (5) and (6) have been made dimensionless by dividing the usual equations by \( mc \omega \) and \( \epsilon \), respectively. Below, these equations are solved by expressing the various quantities in perturbation expansions in terms of \( \epsilon \). Subsequently, terms of like order in \( \epsilon \) will be collected and equated [41]. We use the symbol “\( O_\epsilon \)” to denote “on the order of \( \epsilon \);” that is, \( f = O_\epsilon (\epsilon) \) means \( 0 < \lim_{\epsilon \to 0} f/\epsilon < \infty \), in distinction from \( f = O(\epsilon) \), which is equivalent to \( 0 \leq \lim_{\epsilon \to 0} f/\epsilon < \infty \). Superscripts in parentheses denote the order of the terms.

First, the momentum is expanded as

\[
p = \sum_{i=0}^{\infty} p^{(i)}, \quad p^{(i)} = O_\epsilon (\epsilon^i). \tag{7}
\]

Note that \( p^{(0)} \) would be the momentum in the absence of the standing-wave field, that is, the initial momentum, since \( p \rightarrow p^{(0)} \) as \( a \sim \epsilon \downarrow 0 \). Next, each order \( p^{(i)} \) is decomposed into a slowly varying guiding-center part \( \tilde{p}^{(i)} = (p^{(i)}) \) and a rapidly varying quiver part \( \tilde{p}^{(i)} = p^{(i)} - (p^{(i)}) \). Here, \( \langle \cdot \rangle \) denotes time-averaging on the time scale \( \omega^{-1} \). Upon substitution of this decomposition in the left-hand side of Eq. (5), each term is differentiated with respect to \( \omega t \). This preserves the order of magnitude of the fast quantities \( \tilde{p}^{(i)} \), since these vary on the time scale \( \omega^{-1} \); that is, \( d\tilde{p}^{(i)}/d(\omega t) = O_\epsilon (\epsilon) \). However, from the discussion above, the slow quantities \( \tilde{p}^{(i)} \) and \( \tilde{p}^{(i)} \) vary over longer time scales, so that differentiation increases their magnitude of the fast quantities \( \tilde{p}^{(i)} \) by a factor \( \epsilon \). Therefore the transverse guiding-center part changes on a third, still longer time scale \( \epsilon^2 \omega^{-1} \).

Next, expansion (7) is substituted in the reciprocal Lorentz factor \( \gamma^{-1} \). Extracting the zeroth-order part \( \gamma^{(0)} = \sqrt{1 + (p^{(0)})^2/(mc)^2} \), this gives

\[
\gamma^{-1} = \frac{1}{\gamma^{(0)} \sqrt{1 + [mc\gamma^{(0)}]^{-2}[2p^{(0)} \cdot p - p^{(1)} + O(\epsilon^2)]}}.
\]

\[
\gamma^{-1} = \frac{1}{\gamma^{(0)}} - \frac{p^{(0)} \cdot p^{(1)}}{(mc)^2[\gamma^{(0)}]^{-2}} + O(\epsilon^2).
\]

The quantity \( 1/\gamma^{(1)} \) is the first-order time-dependent variation of the reciprocal Lorentz factor with respect to the constant value \( 1/\gamma^{(0)} \). As shown below, this variation leads to an additional contribution in the final ponderomotive force expression.

Finally, below it is required to take the time average of expressions involving powers of \( A \) or its derivatives. These time averages need to be taken along the trajectory of the particle; that is, in the average \( \langle A[x(t), t] \rangle \), the vector potential is to be evaluated at \( x = x(t) \). To bring out this position dependence explicitly, the position is also expanded in a perturbation expansion in terms of \( \epsilon \):

\[
x = \sum_{i=0}^{\infty} x^{(i)}; \quad \frac{dx^{(i)}}{dt} = O_\epsilon (p^{(i)}) = O_\epsilon (\epsilon^i). \tag{9}
\]

Again, each order \( x^{(i)} \) is decomposed into a slowly varying guiding-center part \( \tilde{x}^{(i)} = x^{(i)} \) and a rapidly varying quiver part \( \tilde{x}^{(i)} = x^{(i)} - (x^{(i)}) \). Then the vector potential can be expanded in a Taylor series around \( x = \sum x^{(i)} = \tilde{x} \):

\[
A[x(t), t] = \tilde{A} + \tilde{z}^{(1)}(\epsilon) \frac{d\tilde{A}}{d\tilde{z}} + O(\epsilon^3). \tag{10}
\]

Here and in the remainder of the paper, an overbar on the vector potential denotes evaluation at the guiding-center position; that is, \( \tilde{A} \equiv A[\tilde{x}, t] \). In writing the series in Eq. (10), it has been anticipated that \( \tilde{x}^{(0)} = 0 \), as shown below.
C. Order-by-order solution of equations of motion

Substituting expansions (7)–(10) in equations of motion (5) and (6), and collecting terms of equal order in \( \epsilon \) results in two equations at each order of \( \epsilon \). These order equations are listed in Appendix A. Order-by-order solution, balancing in each equation the averaged parts and the oscillating parts separately, yields the zeroth-order quantities

\[
\begin{align*}
\overline{p}^{(0)} &= p_0, \\
\overline{p}^{(0)} &= 0, \\
\frac{d\overline{\mathbf{x}}^{(0)}}{dt} &= \frac{p_0}{m} \\
\overline{\mathbf{x}}^{(0)} &= 0,
\end{align*}
\]

in which \( p_0 \) is the initial momentum. As expected, at zeroth order (that is, in the limit \( a \sim \epsilon \downarrow 0 \), where both the field strength and the field inhomogeneity are 0), the motion is equal to what it would be if the EM field were absent. For first-order quantities, it is found that

\[
\begin{align*}
\overline{p}^{(1)}_z &= -eA_z, \\
\frac{d\overline{p}^{(1)}_z}{dt} &= \frac{e}{m} \frac{\partial A_z}{\partial z} \cdot p_{0\perp}, \\
\frac{d\overline{z}^{(1)}}{dt} &= \frac{1}{m} \left[ p_0 - \frac{eA_z}{mc} \right] p_{0\perp} - \left( \frac{eA_z}{mc} p_{0\perp} - \frac{p_0}{m} \right), \\
\frac{d\overline{\mathbf{z}}^{(1)}}{dt} &= \frac{p^{(1)}_\perp}{m}.
\end{align*}
\]

Equation (15) expresses the well-known result that, in an oscillating EM field, at lowest order the quiver momentum balances the vector potential, such that the canonical momentum \( p + eA \) is conserved. Equation (18) is used in Sec. III for the description of the guiding-center motion. Equations (16) and (17) describe the quiver motion in the direction normal to the plane of polarization of the standing wave, which is the direction of the strong field gradient. This is the magnetic quiver motion described in Sec. I. Time differentiation of Eq. (17) and substitution of Eq. (16) yield

\[
\frac{d^2\overline{z}^{(1)}}{dt^2} = \frac{e}{m} \left[ \left( \beta_0^2 - \beta_0 \right) \frac{\partial A_\perp}{\partial z} + \beta_0 \frac{dA_\perp}{dz} \right] \cdot p_{0\perp},
\]

where \( \beta_0 = p_0 / (mc\gamma_0) \) is the initial velocity divided by \( c \). We now restrict to the situation where \( \beta_0 \) is sufficiently small that the second term in brackets in Eq. (19) is negligible, which is the case if \( \beta_0 \ll 1 \). Then, in addition, using Eq. (3) the full time derivative may be written \( d^2/\gamma dt^2 = (\partial/\gamma dt + v \cdot \nabla)^2 = -\sigma^2 + O(\epsilon, \delta, \beta_0) \), so that double time integration yields

\[
\overline{z}^{(1)} = -\frac{e}{m} \frac{\partial A_\perp}{\partial z} \cdot p_{0\perp} \left[ 1 + O(\delta, \beta_0) \right].
\]

The second bracketed term expresses the error introduced by the integration. Equation (20) clearly shows that the amplitude of the magnetic quiver motion is polarization dependent: if \( p_{0\perp} \) is parallel to \( A_\perp \), this amplitude may be substantial, while for \( p_{0\perp} \) perpendicular to \( A_\perp \), it vanishes at first order. This reinstates the argument made in Sec. I: if \( p_{0\perp} \parallel A_\perp \), the momentum is largely perpendicular to the magnetic field, resulting in a substantial magnetic force and quiver amplitude. Conversely, if \( p_{0\perp} \perp A_\perp \), the momentum is largely parallel to the magnetic field, with vanishing magnetic force and quiver amplitude.

As the final step, we substitute Eqs. (11)–(15) and (20) in the right-hand sides of the remaining order equations, (A7) and (A8), and take the time average of these equations. Then the left-hand side reduces to the rate of change of the first-order guiding-center momentum, \( dp^{(1)} / dt \). The right-hand sides reduce to a single gradient:

\[
\frac{dp^{(1)}_z}{dt} \approx -\frac{e^2}{2m\gamma_0} \\
\cdot \left[ \nabla \left( \frac{A_\perp^2}{2} - (\beta_{0\perp} \cdot A_\perp)^2 \right) + \left[ \frac{\partial (\beta_{0\perp} \cdot A_\perp)}{\partial (kz)} \right]^2 \right].
\]

This is the polarization-dependent ponderomotive force in a nonrelativistic standing wave for a particle with \( \beta_{0\perp} \ll 1 \); it is the main result of this paper. The approximate sign expresses a relative error of the order of \( \beta_{0\perp} \). In the limit \( \beta_{0\perp} \rightarrow 0 \), Eq. (21) reduces to the well-known polarization-independent ponderomotive force, Eq. (1), with the relativistic factor \( \gamma_0 \approx 1 + p^2/(mc)^2 + a^2/2 \) included in the denominator. For \( \beta_{0\perp} \neq 0 \), however, the two polarization-dependent terms in Eq. (21) become significant.

The term in Eq. (21) proportional to \( [\partial (\beta_{0\perp} \cdot A_\perp)]/\partial (kz) \) originates from including the magnetic quiver motion, Eq. (20), in the Taylor expansion of the vector potential, Eq. (10). It accounts for the fact that the \( z \) position oscillates in phase with the temporal oscillation of the field. Therefore, when a field gradient in the \( z \) direction is present, the particle systematically samples higher fields at selected phases of the electric quiver motion in the direction of \( A_\perp \) and lower fields at other phases. The induced nonzero average force is negligible in most applications of the ponderomotive force. But in a standing wave the derivative \( \partial/\partial (kz) \) is of order unity, so that this term is comparable to the other terms in Eq. (21).

The origin of the term in Eq. (21) proportional to \( (\beta_{0\perp} \cdot A_\perp)^2 \) can be traced back to including the first-order term \( 1/\gamma_0 \) in the expansion of the Lorentz factor, Eq. (8), rather than approximating \( \gamma \approx \gamma_0 \) throughout.

D. Time-averaged Hamiltonian

The ponderomotive force on a charged particle in an EM field can also be derived using a Hamiltonian description, showing that this force is essentially the gradient of the time-averaged Hamiltonian of the particle [42]. Indeed, the right-hand side of Eq. (21) can be written as such. To see this, note that combination of Eqs. (15), (16), and (20) gives

\[
\frac{dp^{(1)}_z}{dt} \approx -\nabla \left( \gamma_0 mc^2 + \frac{|p^{(1)}|^2}{2\gamma_0 m} - \frac{|p_0 \cdot p^{(1)}|^2}{2\gamma_0 mc^2} \right),
\]

where the term \( \nabla (\gamma_0 mc^2) \equiv 0 \) has been added for convenience. The quantity in angle braces is just the average of the Hamiltonian

\[
H = \sqrt{m^2c^4 + p^2c^2},
\]

in which the momentum \( p \) and the square root have been expanded up to second order in \( \epsilon \).
III. WIGGLING MOTION IN A STANDING WAVE

The last term of the z component in Eq. (21) may be rewritten by performing the z differentiations and using Eq. (3), after which Eq. (21) becomes

$$\frac{d^2k_{\perp}}{d(\omega t)^2} \approx -\frac{e^2}{2m\gamma(0)^2} \frac{\partial}{\partial z}(k_{\perp}^2 - 2(\beta_{0z} \cdot \mathbf{A}_{\perp})^2).$$  \hspace{1cm} (22)

We now evaluate Eq. (22) for the linearly and circularly polarized standing waves considered by Pokrovsky and Kaplan.

A. Linear polarization

Let the standing wave be produced by two counterpropagating plane waves of equal amplitude and frequency that are collinearly polarized in the direction $e_x = e_x \cos \phi + e_y \sin \phi$, as indicated in Fig. 1 by the dashed arrow. Then the vector potential is $A = A_0 e_p \cos k z \sin \omega t$. Suppose that a charged particle enters the standing wave parallel to the $x$ axis with initial velocity $\beta_0$. Differentiating Eq. (18), and substituting Eq. (22), gives the equation of motion for the guiding center of the particle in the $z$ direction:

$$\frac{d^2k_{\perp}}{d(\omega t)^2} + \frac{e^2}{2\gamma(0)^2} \left[ 1 - 2\beta_0^2 \cos^2 \phi \right] \sin(2k_{\perp}) \approx 0,$$  \hspace{1cm} (23)

in which $a_0 = eA_0/(mc)$. Equation (23) shows that the guiding center makes pendulum-like oscillations in the $z$ direction (it wiggles in the ponderomotive potential), with equilibrium points at $k_{\perp} = n\pi/2$ and low-amplitude frequency $\Omega = a_0\sqrt{1 - 2\beta_0^2 \cos^2 \phi}/(\sqrt{2}\gamma(0))$.

First, consider polarization parallel to the initial velocity, that is, $\cos \phi = 1$. Then Eq. (23) reduces to Eq. (40) in Ref. [38] after rewriting $\beta_0 = p_0/(mc\gamma(0))$. As noted in Ref. [38], the most striking feature of this configuration is that the guiding-center oscillations vanish for $\beta_0 = 1/\sqrt{2}$, in terms of the guiding-center motion, the standing wave is invisible to the particle for this value of initial velocity. When $\beta_0$ is increased above $1/\sqrt{2}$, the stable and unstable equilibrium points in Eq. (23) reverse their positions; that is, the ponderomotive force changes direction toward high field regions. This is the relativistic reversal described in Ref. [38].

For polarization perpendicular to the initial velocity of the particle, $\cos \phi = 0$. Then Eq. (23) becomes identical to Eq. (52) in Ref. [38]. For this polarization, the magnitude of the ponderomotive force and the wiggling frequency are independent of the initial velocity, and the relativistic reversal effect is absent.

B. Circular polarization

If the standing wave is produced by two counterpropagating plane waves that are circularly polarized with opposite helicities, the vector potential is equal to $A = A_0 \cos k z (e_x \sin \omega t + e_y \cos \omega t)$. Thus the standing wave has equally spaced nodes and antinodes along the $z$ axis, while locally the field direction rotates around the $z$ axis with time. Again combining Eqs. (18) and (22) and substituting the vector potential now gives

$$\frac{d^2k_{\perp}}{d(\omega t)^2} - \frac{a_0^2}{2\gamma(0)^2} \sin(2k_{\perp}) \approx 0,$$  \hspace{1cm} (24)

which equals Eq. (56) in Ref. [38] taking into account that the field amplitude used there is $A_0/\sqrt{2}$. Thus, in this configuration the equilibrium points are the same as in the linearly polarized case described by Eq. (23) for $\cos \phi = 0$, although the magnitude of the ponderomotive force is a factor of $[\gamma(0)]^2$ weaker.

For counterpropagating circularly polarized waves of equal helicity, the vector potential reads $A_0 \sin \omega t (e_x \sin k z + e_y \cos k z)$. In this case, the standing wave is polarized in a helix along the $z$ axis. Equations (18) and (22) yield

$$\frac{d^2k_{\perp}}{d(\omega t)^2} + \frac{e^2}{2\gamma(0)^2} \sin(2k_{\perp}) \approx 0,$$  \hspace{1cm} (25)

which is the same as Eq. (62) in Ref. [38] after rewriting $\beta_0 = p_0/(mc\gamma(0))$ and taking into account the different definition of the field amplitude. As noted in Ref. [38], the field intensity is homogeneous along the $z$ axis, so that Eq. (1) predicts zero ponderomotive force. However, the modified ponderomotive force expression, Eq. (21), shows that this force is nonzero, so that Eq. (25) still yields wiggling motion of the guiding center.

In summary, all the equations of motion (23)–(25), can be derived from a single ponderomotive force expression, Eq. (21). We have therefore generalized the results of Kaplan and Pokrovsky, who started from the Lorentz force equation for each individual case. Moreover, Eq. (21) also gives the transverse component of ponderomotive force, which was not considered in Ref. [38].

IV. COMPARISON WITH SIMULATIONS

In this section, Eq. (21) is tested against numerical simulations of electron trajectories in a realistic, nonidealized standing-wave field. We have used the GPT code, which uses an embedded fifth-order Runge-Kutta method with adaptive step-size control [43]. For comparison with the simulations, Eq. (21) is needed in terms of the electric field rather than the vector potential. From Eq. (3), the potential is approximately harmonic in time, so that $(\mathbf{E}_{\perp})_z \approx \omega^{-2}(\mathbf{E}_{\perp})_z$. Therefore $\mathbf{E}_{\perp}$ may be effectively replaced by $\mathbf{E}_{\perp} \omega$ in Eq. (21).

We consider again the configuration shown in Fig. 1, this time with two identical Gaussian laser beams in the fundamental mode that counterpropagate along the $z$ axis, have a central wavelength $\lambda = 800$ nm and a peak intensity $I_0 = 2.0 \times 10^{14}$ W/cm$^2$, and are focused in a common waist of size $w_0 = 12.5$ $\mu$m at $z = 0$. For these parameters, $\epsilon \sim (k w_0)^{-1} = 0.01$ and $a = 0.01$ so that the theory in Sec. II is valid. The beams are assumed to be pulsed with a Gaussian pulse shape of length $\sigma = 85$ fs and timed such that the centers of the pulses coincide at $z = 0$ at time $t = 0$. This configuration yields a standing wave around the origin with length $\sim \sigma$ and width $\sim w_0$, and with a lifetime of $\sim \sigma$ around $t = 0$. Furthermore, we assume that the pulses are collinearly polarized in the direction $e_p = e_x \cos \phi + e_y \sin \phi$, similar to the case considered in Sec. III A. In each of the following cases, the trajectory is calculated for an electron entering with initial velocity $\beta_0 = \beta_0 e_x$ for $t \ll -\sigma$ and with the initial position chosen such that, at $t = 0$, the position of the electron would be $(x, y, z) = (0, w_0/2, \lambda / 8)$ in the absence of the laser fields.
the electron meets the standing wave close to the origin, has
interaction with it for some time \( \sim \sigma \), and leaves the interaction
region in a deflected direction. To maximize the ponderomotive
effects, the initial position has been chosen such that the
electron samples the highest available field gradients both in
the \( z \) direction and in the perpendicular direction.

In Figs. 2 and 3, the trajectory of the electron is shown as
viewed from the negative \( y \) axis, for several values of \( \beta_0 \)
and for two polarization directions. The trajectories are calculated
using three methods:

1. (GPT) numerical integration of the exact equations of
   motion using the GPT code (solid lines);
2. (OLD) calculation according to the classical pondero-
motive force, Eq. (1) (dashed lines); and
3. (NEW) calculation according to our polarization-
dependent ponderomotive force, Eq. (21), with \( \mathbf{A}_\bot \)
   replaced by \( \mathbf{E}_\bot / \omega \) (dash-dotted lines).

Standard paraxial field expressions have been used [18];
these are listed in Appendix C for reference.

Figure 2 shows the configuration in which the polarization
is perpendicular to the initial velocity. In this case the
polarization-dependent terms in Eq. (21) vanish, and Eq. (21)
reduces to the classical expression, Eq. (1). Indeed, for all
three initial velocities the three descriptions yield identical
trajectories to within the width of the lines, showing that
the classical expression gives an excellent description of the
averaged motion of the electron. The usual behavior can be
seen, in which the electron is deflected toward low-intensity
regions.

In Fig. 3, however, the situation is very different. Here,
the polarization is parallel to the initial velocity, so that the
polarization-dependent terms in Eq. (21) become important.
Because of this, in Fig. 3(a) the magnitude of the ponderomotive
force is smaller than in the corresponding case in Fig. 2
(\( \beta_0 = 0.5 \)), the ponderomotive force vanishes in Fig. 3(b),
and it even changes direction toward the high-intensity region
in Fig. 3(c). In all of these cases, the resulting trajectories
are excellently predicted by Eq. (21). Figure 3 demonstrates
the relativistic reversal described in Ref. [38] and Sec. III A.
Meanwhile, Eq. (1) is insensitive to the polarization direction,
so that the trajectories (dashed lines) are incorrectly predicted
to be identical to the corresponding cases in Fig. 2.

The insets in Fig. 3 are closeups of the solid lines, showing
that these actually consist of the smooth, time-averaged
trajectory predicted by Eq. (21) (dash-dotted line) and the GPT
trajectory. The latter also contains the quiver motion, which
has components in the \( x \) direction [Eq. (15)] and the \( z \) direction
[Eq. (20)] and is therefore visible in the plane of drawing in
Fig. 3. The frequency of this quiver motion is \( \omega \), while the
electron moves forward with a velocity \( \beta_x \approx \beta_0 \), so that the
spatial period of the quiver motion is \( 1/(\beta_0 \lambda) \). This has been
made visible by scaling the \( x \) axes in the insets. In Fig. 2, the
quiver motion would not be visible in a closeup because it is
perpendicular to the plane of drawing, since $\mathbf{A} \cdot p_\perp = 0$ in Eq. (20) and the $z$ component of the quiver motion vanishes.

To show some ponderomotive effects in the perpendicular direction as well, in Fig. 4 a sideview is given of the trajectories in Fig. 3(a), as seen from the positive $z$ axis. Note from the vertical scale that in this direction the deflection of the electron is very small due to the very small field gradient. Nevertheless, again it is clear from the figure that Eq. (21) accurately predicts the electron trajectory, contrary to the classical expression. Thus Eq. (21) gives a precise description of the three-dimensional electron trajectory in a realistic, nonidealized field configuration. We have also repeated the simulations for combinations of $\epsilon$ and $a$ other than $a \sim \epsilon$. Whenever both $\epsilon$ and $a$ are less than about 0.1, consistent with the assumptions $a, \epsilon \ll 1$ made in Sec. II, we find the same level of agreement with the GPT results.

V. CONCLUSION

The classical polarization-independent ponderomotive force expression is commonly used to describe the time-averaged motion of a charged particle in an inhomogeneous oscillating EM field. It is generally assumed that this is an accurate description, at least for nonrelativistic field intensities. However, we have shown that this is not always true. If the field configuration possesses a direction in which the field changes on the scale of the wavelength, that is, in a standing wave, and, in addition, the charged particle is relativistic, the ponderomotive force is modified. In particular, it becomes dependent on the polarization of the field. Because of this, the ponderomotive force may even vanish, or change its direction toward high field regions, as found earlier by Kaplan and Pokrovsky [37]. We have derived the modified ponderomotive force expression for these configurations, which is of gradient form like the classical expression. Comparison with simulations in the case of a realistic, nonidealized, three-dimensional field configuration confirmed the general validity of the analytical results.

The modifications of the ponderomotive force derived in this paper may have important implications for applications that involve the ponderomotive interaction of relativistic charged particles and standing EM waves. For example, in the electron bunch length measurement based on ponderomotive scattering of the electrons by a standing wave [19], the polarization of the wave is essential for an optimal design of the experimental setup. In the proposed x-ray-free electron laser relying on the wiggling motion of electrons induced by the ponderomotive force in a standing wave [21], the frequency of wiggling and hence that of the stimulated radiation directly depend on the polarization. Experimental tests involving the controlled scattering of electrons by a standing wave have confirmed the classical ponderomotive force expression [28] and Kapitza-Dirac diffraction [29] using nonrelativistic electrons. It would be very interesting to extend these experiments to relativistic electrons to test the polarization-dependent ponderomotive force, Eq. (21), in the classical limit and to study the polarization dependence of Kapitza-Dirac diffraction.

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APPENDIX A: ORDER EQUATIONS

Substituting expansions (7)–(10) in equations of motion (5) and (6), and collecting terms of equal order in $\epsilon$, results in the following order equations. The components of the vector potential and the spatial derivatives have been treated as $eA_{\perp}/(mc) = O_{\epsilon}(\epsilon)$, $eA_{\perp}/(mc) = O_{\epsilon}(\epsilon^2)$, $\lambda V_{\perp} = O_{\epsilon}(\epsilon)$, and $\lambda \partial/\partial z = O_{\epsilon}(1)$, consistent with Eq. (2) and the assumption $a \sim \epsilon$. For reasons of clarity, the equations are displayed in dimensional form.

\begin{align}
O_{\epsilon}(1): \\
\frac{d\vec{p}^{(0)}}{dt} & = 0; \\
\frac{d(\vec{x} + \vec{x}^{(0)})}{dt} & = \frac{(\vec{p} + \vec{p}^{(0)}_{\perp})}{m\gamma^{(0)}}. \\
O_{\epsilon}(\epsilon): \\
\frac{d\vec{p}^{(1)}_{\perp}}{dt} + e\frac{dA_{\perp}}{dt} & = 0; \\
\frac{d[P^{(0)}_{\perp} + P^{(1)}_{\perp}]}{dt} & = e\frac{\partial A_{\perp}}{\partial \gamma^{(0)}}(\vec{p} + \vec{p}^{(0)}_{\perp});
\end{align}
\[ \frac{d(\bar{x} + \bar{z})^{(1)}}{dt} = \frac{(\bar{p} + \bar{p})^{(1)}}{m\gamma^{(0)}} = \frac{(\bar{p} + \bar{p})^{(0)}(\bar{p} + \bar{p})^{(1)}(\bar{p} + \bar{p})^{(0)}}{m\gamma^{(0)}} \tag{A5} \]

Choosing \( \epsilon \):
\[ \frac{d[\bar{p}_+^{(0)} + \bar{p}_+^{(1)}]}{dt} = \frac{e}{m\gamma^{(0)}}(\nabla_\perp \mathbf{A}_\perp) \cdot (\bar{p} + \bar{p})^{(0)} \tag{A6} \]

\[ \frac{d[\bar{p}_z^{(1)} + \bar{p}_z^{(2)}]}{dt} + e\frac{d\mathbf{A}_z}{dt} = \frac{e}{m\gamma^{(0)}} \left\{ \bar{z}^{(1)}(\nabla_\perp \frac{\partial \mathbf{A}_\perp}{\partial z}) \cdot (\bar{p} + \bar{p})^{(0)} + \frac{\partial \mathbf{A}_\perp}{\partial z} \cdot (\bar{p} + \bar{p})^{(1)} + \frac{\partial \mathbf{A}_\perp}{\partial z}(\bar{p} + \bar{p})^{(0)} \right\} \tag{A7} \]

Choosing \( \epsilon^3 \):
\[ \frac{d[\bar{p}_+^{(1)} + \bar{p}_+^{(3)}]}{dt} = \frac{e}{m\gamma^{(0)}} \left\{ \bar{z}^{(1)}(\nabla_\perp \frac{\partial \mathbf{A}_\perp}{\partial z}) \cdot (\bar{p} + \bar{p})^{(0)} + (\nabla_\perp \mathbf{A}_\perp) \cdot (\bar{p} + \bar{p})^{(1)} + (\nabla_\perp \mathbf{A}_\perp)(\bar{p} + \bar{p})^{(0)} \right\} \tag{A8} \]

### APPENDIX B: GENERAL FIELD STRENGTH

The ponderomotive force, Eq. (21), has been derived under the assumption that \( a \sim \epsilon \). Since this expression is the result of balancing terms of equal order in the equations of motion, one might expect that it would be affected by changing the order of magnitude of the vector potential to, for instance, \( a \sim \epsilon^2 \). This is not the case, however. We only give a sketch of the generalized derivation for arbitrary \( a \ll 1 \).

Repeating first the order expansion method for the case \( a \sim \epsilon^n \), \( n \geq 1 \), it is not difficult to find the lowest order slowly varying term and the lowest order rapidly varying term of the expansions of \( p \) and \( 1/\gamma \). Also, the first two terms in the expansion of \( A \) follow straightforwardly. The right-hand side of the equation of motion (5), is then formed by factoring out the product of these three expansions. Taking the time average of the result, it is found that the lowest order terms that are nonzero on average are precisely those that form the ponderomotive force given by Eq. (21). In terms of the corresponding set of order equations analogous to those in Appendix A, the first \( 2n - 1 \) orders of the momentum equations yield vanishing right-hand sides upon averaging, while the \( 2n \)th and \( (2n + 1) \)th orders evaluate to the \( z \) component and the perpendicular component in Eq. (21), respectively.

Also, the opposite situation, in which \( 1 \gg a \sim \epsilon^{1/n} \), \( n \geq 1 \), is possible. Since in this case factors of \( A \) in the equations of motion lead to terms of fractional order in \( \epsilon \), it is appropriate to expand all quantities in power series in terms of \( \epsilon^{1/n} \) rather than \( \epsilon \). This is effected by using the same power series expansions as before, with the understanding that \( p^{(i)} = O_i(\epsilon^{i/n}) \) rather than \( p^{(i)} = O_i(\epsilon^i) \), for example. Except for this modification, the derivation of the ponderomotive force is analogous to that for the case \( a \sim \epsilon^n \) considered above, and again, Eq. (21) is found. Thus Eq. (21) is valid for arbitrary \( a \ll 1 \).

### APPENDIX C: FIELD EXPRESSIONS USED IN NUMERICAL CALCULATIONS

The solid lines in Figs. 2–4 have been calculated by numerical integration of the equations of motion,
\[ \frac{dp}{dt} = e \left( E + \frac{1}{\gamma} p \times B \right) \]
\[ \frac{dx}{dt} = \frac{p}{m\gamma} \]
using for \( E \) and \( B \) the following paraxial Gaussian beam fields [18]. For polarization in the \( x \) direction \( (\phi = 0) \),
\[ E = E_+ \exp \left[ -\left( \frac{z + ct}{4(\sigma z)^2} \right) \right] + E_- \exp \left[ -\left( \frac{z - ct}{4(\sigma z)^2} \right) \right] \]
\[ E_\pm = E_0 \frac{w_0}{w} \left( e_x \cos \psi_\pm \pm \frac{xw_0}{z_R w} e_x \sin \chi_\pm \right) \]
\[ B = B_+ \exp \left[ -\left( \frac{z + ct}{4(\sigma z)^2} \right) \right] + B_- \exp \left[ -\left( \frac{z - ct}{4(\sigma z)^2} \right) \right] \]
\[ B_\pm = \frac{E_0 w_0}{c} \frac{w}{w} \left( \mp e_y \cos \psi_\pm - \frac{yw_0}{z_R w} e_z \sin \chi_\pm \right) \]
In which \( E_0 = \sqrt{2I_0/(\varepsilon_0 c)} \) is the peak electric field amplitude, \( w = w_0 \sqrt{1 + z^2/z_R^2} \) is the beam waist, \( z_R = kw_0^2/2 \) is the Rayleigh length, and the Gouy phases are
\[ \psi_\pm = \omega t \pm \left( k_z - \arctan \frac{z}{z_R} + \frac{z^2 + y^2}{z_R w^2} \right) \]
\[ \chi_\pm = \omega t \pm \left( k_z - 2 \arctan \frac{z}{z_R} + \frac{z^2 + y^2}{z_R w^2} \right) \]
The dashed lines in Figs. 2–4 have been calculated according to Eq. (1) with \( I(x,t) = \varepsilon_0 c (\vec{E}_z^2) \) using Eq. (C1) for \( \vec{E}_z \). The
dash-dotted lines have been calculated according to Eq. (21) with $\mathbf{A}_⊥$ replaced by $\mathbf{E}_⊥/\phi$. For polarization in the $y$ direction ($\phi = \pi/2$), replace $x \to y$ and $y \to -x$ in the expressions above.