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Citation for published version (APA):

DOI:
10.1017/jfm.2018.751

Document status and date:
Published: 25/12/2018

Document Version:
Accepted manuscript including changes made at the peer-review stage

Please check the document version of this publication:

- A submitted manuscript is the version of the article upon submission and before peer-review. There can be important differences between the submitted version and the official published version of record. People interested in the research are advised to contact the author for the final version of the publication, or visit the DOI to the publisher's website.
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Download date: 04. Aug. 2019
Velocity and acceleration statistics in rapidly rotating Rayleigh–Bénard convection

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(Received xx; revised xx; accepted xx)

Background rotation causes different flow structures and heat transfer efficiencies in Rayleigh–Bénard convection (RBC). Three main regimes are known: rotation-unaffected, rotation-affected and rotation-dominated. It has been shown that the transition between rotation-unaffected and rotation-affected regimes is driven by the boundary layers. However, the physics behind the transition between rotation-affected and rotation-dominated regimes are still unresolved. In this study, we employ the experimentally obtained Lagrangian velocity and acceleration statistics of neutrally buoyant immersed particles to study the rotation-affected and rotation-dominated regimes and the transition between them. We have found that the transition to the rotation-dominated regime coincides with three phenomena; suppressed vertical motions, strong penetration of vortical plumes deep into the bulk and reduced interaction of vortical plumes with their surroundings. The first two phenomena are used as confirmations for the available hypotheses on the transition to the rotation-dominated regime while the last phenomenon is a new argument to describe the regime transition. These findings allow us to better understand the rotation-dominated regime and the transition to this regime.

Key words: Rotating Rayleigh–Bénard convection, Lagrangian analysis, Rotating turbulent flows, Velocity and acceleration statistics.

1. Introduction

Rotating Rayleigh–Bénard convection (RRBC), a layer of fluid heated from below and cooled from above while rotating about a vertical axis perpendicular to the bottom and top plates, is of practical relevance to various flows in nature and industry. Key examples include oceanic and atmospheric currents, the convective outer layer of the Sun, the interior of giant gas planets as well as convective cooling in turbomachinery and chemical vapour deposition on rotating substrates. The wide range of applicability of rotating thermal convection forms the reason why it attracts so much attention and has been extensively studied by laboratory experiments, see e.g. Rossby (1969); Boubnov & Golitsyn (1986); Zhong et al. (1993); Liu & Ecke (1997); Sakai (1997); Vorobieff & Ecke (2002); Kunnen et al. (2008); King et al. (2009); Zhong & Ahlers (2010); Niemela et al. (2010); Kunnen et al. (2010); Weiss & Ahlers (2011a); Kunnen et al. (2014); Ecke & Niemela (2014); Rajaei et al. (2017), numerical simulations, e.g. Julien et al. (1996);

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It is well-known that the application of background rotation on Rayleigh–Bénard convection leads to a significant flow departure from its nonrotating state. Three different regimes can be considered: (I) the rotation-unaffected regime (the flow is characterized by a so-called large scale circulation (LSC), composed of tiny mushroom-shape plumes, and the heat flux remains constant), (II) the rotation-affected regime (rotation-aligned vortical plumes are the main features of the flow and the heat flux increases with increasing background rotation), and (III) the rotation-dominated regime (or geostrophic regime, the flow is turbulent but at the same time dominated by the Coriolis force; the heat flux drops dramatically with increasing background rotation). We refer to these regimes as regimes I, II and III, respectively. These different regimes are plotted in Figure 1. In this figure and its caption, $Ra = \alpha g \Delta TH^3/(\kappa \nu)$ is the Rayleigh number, $Pr = \nu/\kappa$ is the Prandtl number, $Ro = \sqrt{\alpha g \Delta T / H / (2 \Omega)}$ is the Rossby number, $\Gamma = D/H$ is the aspect ratio and $Nu = qH/k\Delta T$ is the Nusselt number. In these dimensionless numbers $g$ is the gravitational acceleration, $H$ is the cell height, $\Delta T$ is the applied temperature difference, $\Omega$ is the rotation rate, $D$ is the cell diameter, $q$ is the mean heat-current density and $k$, $\alpha$, $\nu$ and $\kappa$ are the thermal conductivity, the thermal expansion coefficient, kinematic viscosity and thermal diffusivity of the fluid, respectively. The transition between regimes I and II is strongly dependent on the aspect ratio and to a lesser extent on $Pr$: for $\Gamma = 1$ and $Pr = 4.38$ it happens at $Ro \approx 2.5$ (Weiss et al. 2010; Zhong & Ahlers 2010). The transition between regimes II and III depends strongly on $Pr$ and $Ra$: for similar parameter values to our experiments, it occurs in the vicinity of $Ro = 0.1$ where $Nu(Ro)$ displays a peak value (data from Zhong et al. 2009; Joshi et al. 2016).
Regimes I and II and the transition between them are easily accessible with experiments and numerical simulations, thus they are better understood. On the other hand, the geostrophic regime (regime III) and the transition to this regime are less explored. It is worth mentioning that the geostrophic regime is the most relevant regime for astrophysical and geophysical applications, thus it is of utmost importance to understand this regime. Approaching the geostrophic regime is difficult as it requires both high turbulence intensity and rotationally constrained flow. Nonetheless, few studies have decisively entered the geostrophic regime (Stellmach et al. 2014; Ecke & Niemela 2014; Cheng et al. 2015; Kunnen et al. 2016; Rajaei et al. 2017). Other experiments (Rossby 1969; Zhong et al. 1993; Liu & Ecke 1997; King et al. 2009; Liu & Ecke 2009; Zhong et al. 2009; Weiss et al. 2010; Zhong & Ahlers 2010; Kunnen et al. 2011; Weiss & Ahlers 2011a,b) and DNSs (King et al. 2009; Zhong et al. 2009; Schmitz & Tilgner 2009, 2010; Stevens et al. 2010; Horn & Shishkina 2015) reach regime III but they did not enter deep into this regime. In addition to the experiments and direct numerical simulations (DNS), Julien and coworkers (Sprague et al. 2006; Julien et al. 2012a,b; Rubio et al. 2014) derived a set of asymptotically reduced equations in the limit $Ro \to 0$, with $Ro$ the Rossby number denoting the ratio of the inertial and Coriolis forces. The experimental (Rajaei et al. 2017) and DNS (Stellmach et al. 2014) studies show that these equations are indeed appropriate to study the geostrophic regime of rapidly rotating convection.

As a result of difficulties to explore regime III, our understanding of this regime and the transition to this regime is still limited. Nonetheless, two main mechanisms are proposed for this transition, by King et al. (2009, 2012) and Julien et al. (2012a). King et al. (2009, 2012) suggested that the transition depends on the relative thicknesses of the viscous ($\delta_u$) and thermal ($\delta_T$) boundary layers. They showed that $\delta_T < \delta_u$ in regime II, while $\delta_T > \delta_u$ in regime III, within their range of parameter values. Therefore, they hypothesized that the transition between regimes II and III occurs when the viscous and thermal boundary layer (BL) thicknesses become approximately equal, $\delta_u \simeq \delta_T$. However, a similar transition is also observed from numerical simulations with stress-free boundary conditions and no Ekman layers (Schmitz & Tilgner 2009). On the other hand, Julien et al. (2012a), using the asymptotically reduced equations, suggested that the transition occurs when the bulk is also rotation-dominated and the vortical plumes span throughout the entire domain (columns): the vertical motion in the bulk is then suppressed, resulting in a drop in the heat transfer efficiency. Recently, Kunnen et al. (2016) also performed simulations with stress-free and no-slip boundary conditions showing that the transition does not always coincide with the intersection of the viscous and thermal boundary layer thicknesses, in contrast with King’s hypothesis. The study of Kunnen et al. supports more the point of view by Julien et al. rather than King’s, however, no certain conclusion can be drawn yet (Kunnen et al. 2016).

Up to now, the transition to regime III has thus been evaluated through various parameters. King et al. (2009, 2012) focused on the relative thicknesses of the Ekman and thermal BLs and global heat transfer, measured by laboratory experiments and numerical simulations. Julien et al. (2012a) examined the vertical velocity and temperature fluctuations by simulations of asymptotically reduced equations. Kunnen et al. (2016) evaluated the thermal and Ekman BLs thicknesses, heat transfer efficiency, distribution of dissipation and mean temperature gradients throughout the domain by DNS with no-slip and stress-free boundary conditions. However, the correct diagnostic parameter is not identified yet.

In this paper, we examine regimes II and III and the transition between them from a new perspective: the experimentally obtained Lagrangian velocity and acceleration fluctuations and autocorrelations. The results in regime I are also included, wherever
appropriate, in order to have a comprehensive discussion. We also make the connection with the previous hypotheses when possible. The presented Lagrangian data allows us to examine the flow structures in all three regimes. Furthermore, it opens up new understanding of the physics behind all three regimes, in particular on the rotation-dominated regime and the transition to this regime.

We start with the experimental and numerical parameters used in this investigation in Section 2. Next, the Lagrangian velocity rms values and autocorrelations are discussed in Sections 3 and 4, respectively. In Section 5 the acceleration rms values are presented and discussed. In Section 6 the inertial wave, observed in the velocity and acceleration autocorrelations, is discussed. We summarize our main findings in Section 7.

2. Methods and parameters

Experiments have been performed using the experimental set-up described in detail in Rajaei (2017); Rajaei et al. (2016a, b, 2017). Here we briefly discuss the main features of the set-up. The experimental set-up composes of a cylindrical convection cell and a tracking system. In the following paragraphs, we will first explain the convection cell. Then, the tracking system and post processing of the data are explained.

The cylindrical convection cell, filled with water, is the same as the one used in Rajaei et al. (2016a, b, 2017). The convection cell is 200 mm tall with diameter of 200 mm, resulting in $\Gamma = 1$. A copper plate, connected to a resistance heater, is placed at the bottom. The copper temperature is measured by a thermistor positioned inside a hole at the centre of the copper plate. A cooling chamber seals the cell from above. The coolant and the fluid inside the cell are separated by a thin sapphire plate. The cooling chamber and the copper plate are connected by a Plexiglas cylindrical vessel. In order to avoid the distortion of the illumination from the side, the cylindrical vessel is placed inside a Plexiglas cubic box.

The three-dimensional particle tracking velocimetry (3D-PTV) system used in this study consists of four charge-coupled device (CCD) cameras (MegaPlus ES2020, 1600 x 1200 pixels) positioned above the convection cell. The cameras record the flow field at a frequency of 30 Hz which is adequate to resolve the smallest length and time scales of the flow field. The illumination is provided by four arrays of light-emitting diodes (LEDs). The tracer particles are fluorescent Polyethylene particles (supplied by Cospheric Co., USA) and have a mean density of $1002 \text{ kg/m}^3$ and a diameter of 75-90 $\mu$m. In the current study, we use a particle tracking system based on the system developed at ETH Zürich (Switzerland) (Maas et al. 1993; Malik et al. 1993; Willneff 2003, 2002; Lüthi et al. 2005). All experimental equipment, including the convection cell and the 3D-PTV system, is placed on the rotating table.

The 3D-PTV experiments are performed in two different fashions: high particle concentration (HPC) and low particle concentration (LPC). Depending on the parameters of interest, either HPC or LPC data sets are used. Note that HPC experiments consist of more trajectories at each time step (more data points in space at each time step) but the trajectories are shorter compared to the LPC experiments. The LPC experiments are used for the calculations of root-mean-square values and autocorrelations of the velocity and acceleration. The HPC experiments are only used for the vortex detection technique and the subsequent analysis in Section 4.1. In LPC experiments, an average number of $\sim 500$ randomly distributed particles are tracked at each time step. On the other hand, in the HPC experiments, an average number of $\sim 1600$ ($\sim 2700$) randomly distributed particles are tracked at each time step in the cell centre (close to the top plate).

The experiments are performed at constant $Ra = 1.3 \times 10^9$, $Pr = 6.7$ and $\Gamma = 1$.
while $Ro$ varies between 0.041 and $\infty$ ($\Omega$ between 4.12 rad/s and 0). We analyse the Lagrangian velocity/acceleration root-mean-square (rms) values and autocorrelations. The Lagrangian velocity and acceleration rms values are calculated in the volumes at three different regions, vertically separated and centred at the rotation axis. Region I (centre) covers the height $0.375H < z < 0.625H$ with a volume of $50 \times 50 \times 50$ mm$^3$, Region II ($z = 0.8H$) covers the height $0.75H < z < 0.85H$ with a volume of $50 \times 50 \times 20$ mm$^3$ and region III ($z = 0.975H$) covers the height $0.95H < z < H$ with a volume of $50 \times 50 \times 10$ mm$^3$. However, the Lagrangian velocity and acceleration autocorrelations are calculated based on the data collected in a volume of approximately $80 \times 60 \times 50$ mm$^3(x, y, z)$ for both the centre (covering a volume between $z = 0.375H$ and $z = 0.625H$) and close to the top plate ($z = 0.875H$; covering a volume between $z = 0.75H$ and $z = H$). Note that it is not possible to divide the measurement domain into smaller subvolumes for the Lagrangian autocorrelations since the particles do not reside in a thin layer of fluid for a long time.

3. Lagrangian rms velocity

We start with the presentation of the Lagrangian velocity fluctuations at the $z = 0.5H$ (centre), at $z = 0.8H$ and at $z = 0.975H$. We focus on regimes II and III in rotating RBC and how the velocity fluctuations can contribute to our understanding of these regimes and the transition between them.

In Rajaei et al. (2016b), an (an)isotropy ratio based on the velocity fluctuations has been introduced to treat the large-scale isotropy, $RU = u_{z\text{rms}}/u_{xy\text{rms}}$. Note that the statistics in $x$ and $y$ directions are almost the same, i.e. $u_{xy\text{rms}} = (u_{x\text{rms}} + u_{y\text{rms}})/2 \approx u_{x\text{rms}} \approx u_{y\text{rms}}$. Figure 2(a,b) shows the velocity fluctuations and the inverse of the anisotropy ratio, $1/RU = u_{xy\text{rms}}/u_{z\text{rms}}$, as a function of $Ro$ for three measurement sets for regimes II and III. We plot $1/RU$ instead of $RU$ as it shows the trends more clearly. Regimes I and II of these graphs have been treated in Rajaei et al. (2016b). Here, we limit our discussion mainly to regime III.

The measurements at the cell centre, dark blue symbols in Figure 2(a), show that the horizontal and vertical velocity fluctuations continue to decrease in regime III with decreasing $Ro$, but at faster decay rates. The decay in regimes II and III is a result of the turbulence suppression with increase in background rotation (Chandrasekhar 1961). The large-scale anisotropy becomes even larger with decreasing $Ro$ in regime III, see blue squares in Figure 2(b). The horizontal and vertical rms velocities at the cell centre decrease at a similar rate. However, the ratio $1/RU$ decreases with decreasing $Ro$ because $u_{z\text{rms}}$ is smaller than $u_{x\text{rms}}$. Similar to the results at the centre, at $z = 0.8H$ (cyan symbols) the vertical velocity fluctuation decreases with decreasing $Ro$ due to the turbulence suppression with increase in background rotation. The horizontal component decreases as well, but reduction of the vertical component is stronger than that of the horizontal one. The horizontal velocity fluctuation is governed by two contributing factors: decrease due to turbulence reduction with background rotation and enhancement due to swirling motions, contributed by the vortical plumes in the horizontal plane. Therefore, the horizontal velocity fluctuation decreases at a slower rate than the vertical counterpart. The horizontal and vertical components intersect in regime III: $1/RU$ goes from below one to values slightly above one. At $z = 0.975H$, both horizontal and vertical velocity fluctuations decrease with decreasing $Ro$. However, as explained for $z = 0.8H$, the vertical velocity fluctuation reduces faster compared to its horizontal counterpart due to the enhancement of the horizontal velocities by the vortical plumes. This results in an increase in the ratio $1/RU$ with decreasing $Ro$, as is clear from Figure 2(b).
In conclusion, in regimes II and III, the horizontal velocity fluctuations at $z = 0.975H$ and $z = 0.8H$ decrease at a slower rate compared to their vertical counterparts due to the vortical plumes. In regime III, a faster decay is observed for the vertical component of the velocity throughout the convection cell. This observation supports the hypothesis proposed by Julien et al. (2012a) stating that the strong suppression of the vertical motion results in a drop in heat transfer efficiency in regime III.

4. Lagrangian velocity autocorrelation

The Lagrangian velocity autocorrelation is one of the useful Lagrangian quantities which provides insight into the turbulent diffusivity of a fluid parcel (Taylor 1921). The Lagrangian velocity autocorrelation shows the time period over which the particle velocity remains correlated with the velocity at the previous times. It is expected that the particle velocity remains correlated for a longer time if the flow is coherent. Therefore, the Lagrangian velocity autocorrelation can be used as a measure of flow coherence as well. The Lagrangian velocity autocorrelation is defined as

$$ R_{u_i}^L(\tau) = \frac{\langle u_i(t)u_i(t+\tau) \rangle}{\langle u_i^2(t) \rangle}, $$

(4.1)

with $u_i$ the $i^{th}$ component of the velocity signal, $\tau$ the time lag and $\langle \ldots \rangle$ the ensemble average. Another useful parameter in the evaluation of the coherent structures in turbulent flows is the integral time scale, defined as

$$ T_{u_i}^L = \int_0^\infty R_{u_i}^L(\tau)d\tau, $$

(4.2)

which is a measure for the time a fluid parcel velocity remains correlated providing insight into the large-scale flow coherence.

In the present study the residence time of the particles in the observation domain (estimated as $d_{\text{obs}}/u_{\text{rms}}$ where $d_{\text{obs}}$ is the length/width/depth of the observation volume and $u_{\text{rms}}$ is the velocity rms) is $\sim 25\tau_\eta$, while the autocorrelation coefficient drops below 0.1 at about $15\tau_\eta$ (see Figure 3) for most of the rotation rates. Furthermore, as an a posteriori check $d_{\text{obs}}$ for the length and width is at least 5 times larger than $V_{\text{drift}}T_{u_i}^L$,
where $V_{\text{drift}}$ is the horizontal velocity of the vortical columns; see Rajaei et al. (2017) for the calculation of $V_{\text{drift}}$. They indicate that the bias error induced by the limited volume is small.

For large time lags $\tau$ the autocorrelation values are not perfectly converged due to the shortage of long trajectories. Lack of convergence leads to larger errors in the calculation of the integral time scale. Therefore, we fit an exponential function over the time period for which the velocity autocorrelation shows a clear exponential decay. Exponential fitting is chosen since in homogeneous isotropic turbulence (HIT) the velocity decorrelation at large $\tau$ is expected to be exponential (Mordant et al. 2004a, 2001; Gervais et al. 2007; Del Castello & Clercx 2011); it plays a crucial role in some dispersion models (Sawford 1991). The linear-logarithmic plots (not shown here) confirm the exponential decay of the velocity autocorrelation at large $\tau$. Note that at very small (dissipative scale) $\tau$, in our experiments $\tau \lesssim 1$ s, the velocity autocorrelation shows non-exponential behaviour (the curvature of the velocity autocovariance at $\tau = 0$ is equal to the acceleration variance) (Voth et al. 2002; Mordant et al. 2004b; Sawford 1991). The integral time scale is calculated using a combination of the actual velocity autocorrelation curve for small time lags and the exponentially-fitted curve for large time lags. The integral time scale can also be estimated through the aforementioned exponential fitted curve $e^{-\tau/\tau_0}$, where $\tau_0$ is the integral time scale. However, the latter approach neglects the non-exponential decay for small time lags, so small differences are expected. We shall use the first method (i.e. a combination of the actual velocity autocorrelation curve for small time lags and the exponentially-fitted curve for large time lags) unless otherwise stated.

We start with the data at the centre. Figure 3(a,b) shows the autocorrelation of the velocity in the $xy$ and $z$ directions at the cell centre, respectively. The time lag is nondimensionalised by the local Kolmogorov time scales, given in Table 1 calculated from direct numerical simulations using the same parameter settings, see Alards et al. (2017) and Rajaei (2017) for the details of the numerical method. Note that the Kolmogorov time scale differs depending on the position within the cylinder. In Table 1, we also report the Froude number; a measure for the importance of the centrifugal buoyancy and defined as $Fr = \Omega^2 R / g$ where $R$ is the cell radius. It is worth mentioning that Horn & Aurnou (2018) suggest that centrifugal effects are not important when $Fr < R/H$ (in our case $R/H = 0.5$). The velocity autocorrelations in regime I collapse for both horizontal and vertical components. Figure 4 displays the Lagrangian integral time scales, nondimensionalised by the local Kolmogorov time scale, at the cell centre and at $z = 0.875H$. The horizontal and vertical integral time scales at the cell centre, red circles and blue squares, remain almost constant in regime I as well. As mentioned before, the integral time scale is a measure of the large-scale coherence. Therefore, one expects smaller integral time scales at $Ro \simeq 2.4$, where the LSC breaks down and the flow lacks a dominant large-scale coherent structure. However, Figure 4 does not show a decrease around $Ro \simeq 2.4$. It can be explained by the fact that the LSC is best visible near the top/bottom and side boundaries: the large-scale flow at the cell centre is close to HIT (Sun et al. 2006; Rajaei et al. 2016b) and there is no clear signature of the large-scale coherent structure in the cell centre for $Ro \gg 1$. In regime II, the horizontal and vertical velocity autocorrelations progressively increase with reduction of $Ro$, see Figure 3(a,b). These enhancements are also reflected in the integral time scales, see red circles and blue squares in Figure 4. In regime II, the vortical plumes are already present as the large-scale flow features, resulting in an enhancement of the large-scale coherence. Note that, in contrast to the LSC, the vortical plumes are visible in the large-scale flow in the centre (Rajaei et al. 2016b). In regime III, the vortical plumes form columns and the vertical and horizontal velocity autocorrelations show non-monotonic trends with respect...
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Table 1. Kolmogorov time scale, τ_η (in seconds), as a function of Ro for the measurements at the cell centre and close to the top plate. The dissipation rates are calculated from DNS data (Rajaei et al. 2016b). The dissipation rates from DNS are for slightly different Ro numbers, but we expect that they are still representative for the current experiments given the weak Ro-dependence displayed here.

Figure 3. (a) Horizontal and (b) vertical Lagrangian velocity autocorrelations at the centre and (c) horizontal and (d) vertical Lagrangian velocity autocorrelations close to the top as a function of Ro. The inset in panel (d) is the vertical velocity autocorrelation for Ro = 0.041 at z = 0.875H. The green and red lines are the exponential fits, proportional to e^{-τ/τ_0} with τ_0 = τ_1 for 1 ≲ τ/τ_η ≲ 6 and τ_0 = τ_2 for τ/τ_η ≳ 6, respectively. The dash-dotted black line is located at τ/τ_η = 6.

The flow shows strong horizontal correlation in regime I. The horizontal velocity autocorrelations remain correlated for a long time due to the LSC which is prominently present at this Ro but only reveal little changes in this regime. The horizontal and vertical integral time scales also show non-monotonic behaviours. However, the non-monotonic trends are within the error bars.

We continue with the data at z = 0.875H. Figure 3(c,d) shows the velocity autocorrelations at z = 0.875H for the xy and z directions, respectively. The inset of Figure 3(d) shows R_L^z for Ro = 0.041 with the horizontal axis in logarithmic scale.
Figure 4. Integral time scale as a function of $Ro$ for measurements both at the cell centre and top ($z = 0.875H$). The error is estimated by calculating the integral time scale for different segments of the experimental data. The error is approximately 15%. The open diamonds are estimates of the corresponding filled diamonds, see the text for details.

location, see also the black diamonds in Figure 4. As the LSC vanishes for $Ro \lesssim 2.4$, the horizontal velocity correlation time drops abruptly, see also the black diamonds in Figure 4. The horizontal velocity correlation time at $z = 0.875H$ (black diamonds) in regime II are approximately equal to the horizontal velocity correlation time at the centre (red circles) in regime I, where the LSC is not felt either. As the LSC has vanished, the mean horizontal velocity becomes zero, thus no contribution from the mean velocity is expected in the correlations in regime II. Note that, in contrast to the velocity rms values, the autocorrelation data are calculated for the total velocity signal for $Ro \gtrsim 2.4$ at $z = 0.875H$. For the rest of the data (data at the centre and data for $Ro \lesssim 2.4$ at $z = 0.875H$) the mean velocities are zero. In regime II ($2.4 \lesssim Ro \lesssim 0.1$), the horizontal correlation decreases slightly with decreasing $Ro$. In regime III, $Ro \lesssim 0.1$, the horizontal correlation decreases even further as $Ro$ goes down and its corresponding integral time scale (black diamonds) becomes as small as two times the Kolmogorov time scale ($T_{u_{xy}}/\tau_\eta \sim 2$). The decrease in the horizontal autocorrelation throughout regimes II and III can be explained by knowing that the swirling motion in the horizontal plane near the top plate results in a continuous change of velocity direction in this region. An estimate of the time scale of the decorrelation of the horizontal velocity autocorrelation, due to swirling motions near the top plate, can be calculated by using the horizontal velocity of the fluid parcel and the nominal radius of a vortical plume. If we assume that a fluid parcel follows perfectly a circular path, $u_x$ (or $u_y$) decorrelates when a quarter of the circle circumference is covered. Therefore, an approximation of the decorrelation time is given by $\tau_h = \frac{1}{4} \pi L_c/u_h$, where $u_h$ is the magnitude of the horizontal velocity, $u_h = (u_x^2 + u_y^2)^{1/2}$, and $L_c$ is the so-called critical length, $L_c = 4.8154Ek^{1/3}$ (Chandrasekhar 1961) with $Ek$ the Ekman number defined as $Ek = Ro\sqrt{Pr/Ra}$. The $\tau_h$ values are shown as open diamonds in Figure 4 and they display similar trends as the measured integral time scales. Note that it is expected that the horizontal velocity autocorrelation is affected by the drift velocity of the vortical columns. However, quantifying that effect is rather elusive. One might expect that the integral time scale of the horizontal velocity at the centre follows the behaviour of its counterpart close to the top plate when $Ro$ is small. However, we know that the horizontal swirling velocity becomes smaller as we go down in
the vortex column. Therefore, $u_{xy}$ decorrelates more rapidly (and $T^L_{u_{xy}}$ is smaller) close to the top than near the cell centre, given that we expect the horizontal structure size to be equal close to the top and at the cell centre.

In regime III, an oscillatory behaviour is observed in the horizontal component of the velocity near the top plate, see the tails of the correlations in Figure 3(c). We will treat this intriguing oscillatory behaviour in detail in Section 6 and show that not only Earth’s gravity determines the flow dynamics in our experiment but also Earth’s rotation.

The vertical velocity autocorrelations, Figure 3(d), show no appreciable changes with rotation in regime I. As can be seen from the graph, they show persistent negative values. This is in agreement with our picture with an LSC present where a fluid parcel moves upward on one side and downward on the other side of the cell. Considering the general picture, one expects that the anti-correlation occurs at $0.5d_{obs}/u_{LSC}$, where $d_{obs}$ is the length of the observation domain and $u_{LSC}$ is the fluid velocity in the LSC, calculated based on the LSC Reynolds number (Song & Tong 2010). The anti-correlation starts at $\tau/\tau_\eta \approx 7.0$ ($\tau \approx 7.5$ s), which is compatible with $0.5d_{obs}/u_{LSC} \approx 6$ s. The vertical velocity autocorrelations collapse in regime I, with a rather large anti-correlation; $-0.5$ at $\tau/\tau_\eta = 25$. The LSC vanishes for $Ro \lesssim 2.4$, and the vertical velocity anti-correlation gradually diminishes. For relatively high rotation rates, $Ro \lesssim 0.5$, the anti-correlation in the vertical velocity autocorrelation disappears.

The vertical integral time scales in regime I and part of regime II at $z = 0.875H$ are not reported due to the negative values in the vertical velocity autocorrelations. The vertical integral time scales in regime II and III (pink stars in Figure 4) increase with decreasing $Ro$ which is an indication of enhancement in the flow coherence, which we expect to be due to a stronger vertical coherence of the vortical plumes.

### 4.1. Small Ro: emergence of two time scales

As mentioned before, an exponential decay at large time lags in the velocity autocorrelation is expected; it plays a crucial role in Sawford’s stochastic dispersion model (Sawford 1991). Our experimental data also confirms the exponential decay, proportional to $e^{-\tau/\tau_0}$, in the velocity autocorrelation. As mentioned before, the behaviour of the velocity autocorrelation is far from an exponential decay at the dissipative time lags $\tau/\tau_\eta \lesssim 1$ (Voth et al. 2002; Mordant et al. 2004b; Sawford 1991). However, here we focus on the large time lags $\tau/\tau_\eta \gtrsim 1$ of the velocity autocorrelations.

For all rotation rates at the cell centre and most of the rotation rates at $z = 0.875H$, the slope of the exponential decay is constant for $\tau/\tau_\eta \gtrsim 1$, i.e. a single decay constant $\tau_0$ in the expression $e^{-\tau/\tau_0}$ is observed. However, for the vertical velocity autocorrelation for $Ro \leq 0.19$ at $z = 0.875H$ a second time scale emerges. The inset of Figure 3(d) shows the vertical velocity autocorrelation for $Ro = 0.041$ on a logarithmic scale. The green and red lines are exponential fits, proportional to $e^{-\tau/\tau_i}$, fitted to the part where the autocorrelation shows a clear exponential decay. As can be seen, there exist two time scales $\tau_i$; $\tau_1$ for small time lags ($1 \lesssim \tau/\tau_\eta \lesssim 6$) and $\tau_2$ for large time lags ($\tau/\tau_\eta \gtrsim 6$). The dash-dotted black line is at $\tau = 6$ s, where the change in slope occurs.

We can calculate the two correlation time scales ($\tau_1$ and $\tau_2$) for the vertical velocity autocorrelations plotted in Figure 3(d). The results are plotted in Figure 5 (solid symbols) as a function of $Ro$. The smaller correlation time scales ($\tau_1$), present for smaller $\tau$ and represented by the solid red circles, remain almost unchanged in the regimes II and III while the larger correlation time scales ($\tau_2$), present for larger $\tau$ and represented by solid blue squares, increase with decreasing $Ro$. The pink stars are the vertical integral time scales, taken from Figure 4. As expected, they fall between the solid squares and circles. The open symbols will be discussed later.
In order to study the origin of the two correlation time scales, the flow domain is divided into plume and non-plume regions. The so-called $Q$-criterion (Hunt et al. 1988) is a widely used method for vortex detection. We apply the $Q$-criterion for a three-dimensional flow field: a vortex is defined as a spatial region where

$$Q = \frac{1}{2} \left( \| \overline{\Omega} \|^2 - \| \overline{S} \|^2 \right) > 0,$$

(4.3)

where $\overline{\Omega}$ and $\overline{S}$ are the antisymmetric and symmetric parts of the velocity gradient tensor, respectively. The operator $\| \cdots \|$ represents the Euclidean norm defined as

$$\| A \| = \sqrt{\text{Tr} \left( AA^T \right)}.$$

(4.4)

We have previously applied the $Q$-criterion method on 3D-PTV data (Rajaei et al. 2017).

Detection of the plume and non-plume regions allows for differentiation between velocity autocorrelations inside and outside the plumes. Figure 6 shows the vertical velocity autocorrelations inside and outside the plumes for $Ro = 0.048$. In this figure, the open squares represent the vertical velocity autocorrelation of particles inside the plumes ($Q > 0$). The open circles, on the other hand, show the vertical velocity autocorrelation of the particles in non-plume regions ($Q \leq 0$). The solid and dashed red lines are the fitted lines (proportional to $e^{-\tau/\tau_0}$) through the open squares and circles, respectively. Clearly, the time scale for correlation within the plumes (the inverse slope of the solid red line) is larger than that for outside the plumes (the inverse slope of the dashed red line), indicating larger vertical flow coherence inside the plumes. The time scale values for different rotation rates computed with this method are given in Figure 5 as open symbols.

It is worth mentioning that the solid symbols in Figure 5 are calculated based on the velocity autocorrelation from the low particle concentration (LPC) data set while the plume detection method and consequently the open symbols are calculated based on the data from high particle concentration (HPC) experiments. Convergence for the large
time lags of the HPC data set is questionable, thus the open symbols in Figure 5 have inevitably larger uncertainties and the comparisons should be considered as qualitative; a one by one comparison might not be legitimate. However, the open and solid circles and squares clearly display similar trends.

It can be concluded that $\tau_1$ in the vertical velocity autocorrelation (represented by the green line in the inset of Figure 3(d)), which is the governing time scale at small time lags $1 \lesssim \tau/\tau_\eta \lesssim 6$, is mainly determined by the flow time scale in the non-plume regions. Obviously, there is a contribution from the particles inside the plumes as well, however, this contribution is smaller since the number of particles in the non-plume regions is larger than that inside the plume. As has been shown, the particles in the non-plume regions decorrelate faster than those in the plume regions, e.g. the autocorrelation coefficient at $\tau/\tau_\eta = 6$ is below 0.2 for the non-plume region compared to 0.5 for the plume region in Figure 6. Thus, it is expected that the correlation at large time lags ($\tau/\tau_\eta \gtrsim 6$) is mainly determined by the time scale $\tau_2$ of the plume regions; the non-plume regions are already decorrelated for such large time lags. We interpret the time scale $\tau_2$ of the plume regions as the residence time of the fluid parcel in a vortex, which is most probably closely related to the penetration time of fluid through columns.

Knowing that $\tau_1$ is a characteristic correlation time of the flow outside the plumes, makes it possible to comment on the reason why $\tau_1$ is constant while $\tau_2$ increases with decreasing $Ro$. As $\tau_1$ is the correlation time outside the coherent structures, when $\tau_1$ is nondimensionalised by $\tau_\eta$, one would expect no significant changes with variation in turbulence intensity (i.e. variation in $Ro$). On the other hand, $\tau_2$ is the correlation time of the plumes which is found to vary with $Ro$.

4.2. Concluding remarks

In conclusion, although the first transition (between regime I and II) is reflected in the velocity autocorrelations and the integral time scale (more profoundly near the top plate), the second transition (between regime II and III) is not as clearly reflected. This can be explained by the fact that the former transition is sudden while the latter transition appears to be more gradual. However, the crossover to regime III coincides
with the appearance of an additional correlation time scale in the vertical velocity autocorrelations near the top plate. Although the longer correlation time $\tau_2$ increases with decreasing $Ro$, the shorter time $\tau_1$ remains constant. We have indicated that these two different correlation time scales are associated with the correlations inside and outside the plumes. The emergence of two correlation time scales suggest that the crossover to regime III is caused by the presence of sufficiently strong plumes which interact less with their surroundings. In other words, the fluid exchange between plume and non-plume regions becomes considerably less in regime III. Thus, the plumes are separated from the non-plume regions; each characterized by their own time scale. The flow in regime III has therefore two distinct characteristics, one inside and another outside of the plumes. Further quantitative analysis of this hypothesis using detailed numerical simulations is encouraged.

5. Lagrangian rms acceleration

So far, the discussion was on quantities which are largely determined by the large-scale flow field. In this Section, however, we focus on the Lagrangian acceleration rms values which are governed by comparatively smaller scales. 

Prior to the analysis of the Lagrangian rms acceleration, we first focus on the undesired inertial waves emanating from precession forcing by the interplay of the applied rotation and Earth’s rotation. A discussion about the origin of these waves can be found in Section 6. As mentioned before in Section 4, a wavy behaviour in the horizontal velocity autocorrelation near the top plate is observed. Note that these waves are absent in the results of the numerical simulations (based on the Boussinesq approximation in a rotating frame of reference). The inertial waves have a minor effect on the velocity signal but as expected the effects of the waves are stronger on the acceleration signal. In order to illustrate these effects on the acceleration signal, we look into the Lagrangian acceleration autocorrelations. The Lagrangian acceleration autocorrelation is defined in a similar fashion as the Lagrangian velocity autocorrelation

$$R_{a_i}^L = \frac{\langle a_i(t)a_i(t+\tau)\rangle}{\langle a_i^2(t)\rangle}, \quad (5.1)$$

where $a_i$ is the $i^{th}$ component of the acceleration fluctuation of a fluid parcel. Figure 7 shows the Lagrangian acceleration autocorrelations for horizontal and vertical directions at the cell centre. The effect of the inertial waves are prominent in the Lagrangian acceleration autocorrelation. Therefore, it is expected that the acceleration fluctuations in regime III are also polluted by the presence of the inertial waves. In order to remove the approximate contribution of the inertial waves, we subtract the mean value of the acceleration of all particles in the measurement volume at each time step from the acceleration of each particle at that time step.

For the rms acceleration, similar to Section 3, three different regions are analysed; the cell centre, $z = 0.8H$ and $z = 0.975H$. The Lagrangian acceleration fluctuations and the ratio $RA = a_{xy}^{rms}/a_{z}^{rms}$ are plotted in Figure 8(a,b). In regime III, the acceleration fluctuations without the wave contribution are also calculated; they are plotted as open symbols. For the data at the cell centre and several rotation rates at $z = 0.8H$ and $z = 0.975$ the open symbols are not visible because they almost collapse on the solid symbols. The acceleration fluctuation and its ratio in regimes I and II have been discussed in Rajaei et al. (2016a). Here, however, we focus on the acceleration fluctuation in regime III and the transition from regime II to III.

The horizontal and vertical components of the acceleration at the cell centre, see
blue circles and squares in Figure 8(a), continue to decrease as $Ro$ decreases due to the suppression of turbulence by the background rotation in regime III. The ratio $RA$, which can also be used as a tool to partially evaluate small-scale isotropy, does hardly change while transitioning from regime II to regime III. The differences between open and solid symbols are negligible; they virtually collapse.

The measurements at $z = 0.8H$ show that both horizontal and vertical accelerations decrease hardly in regime II with decreasing $Ro$, resulting in an approximately constant $RA$. However, after the transition to regime III, the horizontal component starts to increase in regime III due to the formation of strong vortical plumes capable of penetrating further into the bulk. The vertical component, on the other hand, keeps decreasing with decreasing $Ro$ in regime III, similar to regime II. As a result, the ratio $RA$ at $z = 0.8H$ increases with decreasing $Ro$ in regime III: a clear indication that the vortical plumes are penetrating further into the interior region.

In regime III close to the top plate ($z = 0.975H$), the horizontal acceleration without the wave contribution shows a slight increase with decrease in $Ro$, similar to that of regime II. The vertical acceleration, on the other hand, continues to decrease in regime III with decreasing $Ro$, but at a higher decay rate compared to regime II. As a result, the ratio $RA$ increases in regime III at an even higher rate. The demarcation between regimes II and III is very well captured by the ratio $RA$.

It is worth mentioning that the decrease in $a^{\text{rms}}_{xy}$ at the cell centre might not directly be related to whether vortical plumes reach the cell centre. Note that the sign of the vertical vorticity will flip as we move down far enough in a vortex column (Portegies et al. 2008; Grooms et al. 2010); the vortical plumes spin down as they approach the centre which effectively reduces $a^{\text{rms}}_{xy}$. The change in sign of vertical vorticity can be explained by the thermal-wind balance (Pedlosky 1979). Together with the decrease of the vertical vorticity, the plumes near the top and bottom walls are also squeezed due to the presence of the boundaries and widened due to conservation of angular momentum.

In conclusion, we can clearly see that the flow dynamics have changed while crossing to regime III near the top plate. If we assume that the vortical plumes interact considerably less in regime III (see the concluding remarks in Section 4.2), a particle inside a plume stays inside the plume, thus experiencing a strong horizontal centripetal and weak vertical accelerations. In regime II, however, fluid parcels (particles) are exchanged between plume and non-plume regions more easily, resulting in a comparatively weaker horizontal centripetal acceleration. Therefore, the increase in the ratio $RA$ in regime III supports the conclusion drawn in Section 4.2: there is less exchange between plume and non-plume regions in regime III.

Furthermore, the horizontal component of the rms acceleration at $z = 0.8H$ starts to
increase while crossing the transition point: the vortical plumes probably penetrate farther into the bulk to form columns. However, at the cell centre the horizontal component of rms acceleration continues to decrease with decreasing Ro in regime III.

6. Oscillatory behaviour

Wavy behaviours are observed in the Lagrangian velocity and acceleration autocorrelations. It is known that in a rotating frame, the Coriolis force promotes waves, so-called inertial waves (Greenspan 1968). One can readily derive the dispersion relation from the inviscid linearised Navier-Stokes equations (Greenspan 1968). Imposing boundary conditions makes the solutions dependent on the geometry; we focus on the cylindrical geometry. There are a variety of studies on inertial waves in a cylindrical geometry under different forcing mechanisms: steady differential rotation (Hart & Kittelman 1996; Lopez & Marques 2010), forced sidewall oscillations (Lopez & Marques 2014), libration (Busse 2010; Le Bars et al. 2015; Noir et al. 2010) and precession (Gans 1970; Kobine 1995; Manasseh 1992; Liao & Zhang 2012; Meunier et al. 2008). The first three conditions are
not relevant for our experimental measurement. We consider the effects of precession due to Earth’s rotation in our experimental measurements. In precessing flows, a cylinder, filled with fluid, rotates about its axis while it precesses about another axis, see Figure 9 for a schematic view of the problem. In this figure $\Omega_P$ is the precession angular velocity, $\Omega$ is the cylinder angular velocity and $\alpha$ is the angle between $\Omega$ and $\Omega_P$.

The study of precessing flows is motivated by its relevance in the field of aerospace (propellant fuel in a spinning spacecraft e.g. Stewartson (1959); Agrawal (1993)) and atmospheric science (tornadoes and hurricanes (Vanyo 2015)). In our experimental set-up, Earth’s rotation can act as the precession angular velocity. The important dimensionless parameters in precessing flows in cylindrical domains are the Ekman number, the Poincaré number and the aspect ratio (Liao & Zhang 2012). The Poincaré number is defined as $\text{Po} = |\Omega_P|/|\Omega|$; in our case $|\Omega_P| = |\Omega_{\text{Earth}}| = 7.292 \times 10^{-5}$ rad/s and $\Omega$ is the rotation rate of the rotating table, $\Omega = \Omega \hat{z}$, resulting in $1.8 \times 10^{-5} \lesssim \text{Po} \lesssim 4.4 \times 10^{-5}$ in regime III. The Poincaré number indicates the strength of precession (Wu & Roberts 2009) while the aspect ratio defines the resonance condition (Gans 1970).

The momentum equation for precessing flow in cylindrical coordinates $(r, \phi, z)$ is given in Liao & Zhang (2012). One can readily write the momentum equation in dimensionless form. We can further simplify the dimensionless momentum equation when $\text{Po} \ll 1$ (weak precession), $\text{Ro} \ll 1$ and $\tilde{u} < 1$ (tilde stands for the dimensionless variable, the velocity is scaled with the free-fall velocity $U = \sqrt{\alpha g \Delta TH}$) (Liao & Zhang 2012)

$$\frac{\partial \tilde{u}}{\partial t} + \tilde{u} \cdot \nabla \tilde{u} + \frac{1}{\text{Ro}} \hat{z} \times \tilde{u} = -\nabla \tilde{P} + \sqrt{\frac{Pr}{Ra}} \tilde{u}^2 - \frac{Po}{4Ra^2} r \sin \alpha \cos \left(\phi + \frac{t}{2\text{Ro}}\right) \hat{z} + \tilde{T} \hat{z}. \quad (6.1)$$

As can be seen from the equation, the third term on the right hand side, the so-called Poincaré forcing term, is the principal forcing term due to precession. To quantify the effects of Poincaré forcing on the flow field in our current experiments, direct numerical simulations with this precession forcing are performed at $Ra = 1.3 \times 10^9$, $Pr = 6.7$, $\text{Ro} = 0.058$, $\sin \alpha = 0.6232$ (Eindhoven latitude = 51.44°) and corresponding $Po = \Omega_P/\Omega = 2.51 \times 10^{-5}$. As in previous sections, tracer particles are tracked and evaluated.

The experimental and corresponding numerical acceleration autocorrelations at the cell centre and at $z = 0.875H$ are plotted in Figure 10. The excellent agreement between experimental and simulation data confirms the role of the Poincaré force emanating from
Earth's rotation. The wave amplitudes are slightly different between experiment and numerical simulation, however, one should consider that the equations are simplified in the numerical simulations. The observed frequency in the acceleration autocorrelations is exactly $\Omega$ in both the experiment and the numerical simulation. For an inviscid weakly precessing fluid, when the frequency of the primary inertial mode is the same as that of the Poincaré force (the frequency of Poincaré force is $\Omega$), a primary resonance happens. As shown by Meunier et al. (2008) the primary resonance for a cylindrical cell with aspect ratio $\Gamma = 1$ happens at a forcing frequency of $0.996\Omega$. Our forcing frequency is $\Omega$, very close to the primary resonance for $\Gamma = 1$. The primary resonances for $\Gamma = 1/2$ and 2 are $0.49\Omega$ and $1.57\Omega$, respectively. Note that these oscillations vanish when the Poincaré term in Equation (6.1) is set to zero (Boussinesq approximation in rotating frame of reference).

The velocity field belonging to the near-resonant inertial waves due to Poincaré forcing can be illustrated by plotting the solution of the wave equations. The wave equations and their solution are given in Meunier et al. (2008). Figures 11(a,b,c) show snapshots of the waves at time $t = 0$. Panel (a) shows the velocity of the waves at three cross sections perpendicular to the rotation axis. The colours and arrows show the vertical and horizontal velocities, respectively. Two vertical cross sections along line $C$ are presented in panels (b) and (c). The colours in panels (b) and (c) are the horizontal velocity along line $C$ and the vertical velocity, respectively. See the figure caption for more details. As can be seen from the graph, the horizontal component of the wave becomes stronger when approaching the top or bottom plates from the centre. The vertical component, on the other hand, is strongest at the cell centre and diminishes when departing from the centre.
Figure 11. Illustration of the dominant inertial wave mode in a $\Gamma = 1$ cylinder. Panel (a) represents three cross sections at $z = 0.85H$, $z = 0.5H$ and $z = 0.15H$ from top to bottom, respectively. The colors indicate the vertical velocity (perpendicular to the cross sections) and arrows represent the horizontal velocity. The horizontal velocity at $z = 0.5H$ is zero. Panel (b) shows a vertical cross section along line $C$, shown in panel (a). The colours indicate the horizontal velocity along line $C$, $u_C$. At $r/H = 0$ the velocity $u_C$ is not defined. Panel (c) shows the vertical velocity in a vertical cross section along line $C$, similar to panel (b). The colour bar and consequently all the velocities are nondimensionalised by their corresponding maximum velocities at each panel.

The same trend has been observed from experiments: strong vertical oscillations at the cell centre and strong horizontal oscillations near the top plate. This pattern rotates about the cylinder axis in an anticyclonic fashion at a rate $\Omega$. Note that these plots are formally for an inviscid flow; corrections for no-slip conditions in a realistic situation are restricted to minor boundary layer corrections (Meunier et al. 2008).

In conclusion, the observed oscillatory behaviour is emanating from precession by Earth’s rotation: these waves are thus unavoidable. However, one can significantly reduce the amplitude of these waves by choosing a proper $\Gamma$ further from a resonant mode (Meunier et al. 2008). Nonetheless, it was unfortunate that $\Gamma = 1$ has a primary resonance at a driving frequency very close to $\Omega$. It is worth pointing out that the good agreement between experimental and numerical overall heat transfer measurements for small $Ro$ for $\Gamma = 1$, reported earlier in literature (Stevens et al. 2009; Zhong et al. 2009), hint that the effects of these inertial waves on the heat transfer efficiency are negligible. However, one should note that the Poincaré forcing term increases with decreasing $Ro$: the Poincaré force dominates over thermal forcing for small enough $Ro$ depending on $\alpha$. The Poincaré force is not limited to the thermally-driven motions: it plays a role whenever the Poincaré force becomes comparable to the other means of forcing. Therefore, experiments in rapidly rotating flows should be done with special care: the geometry and its aspect ratio should be chosen accordingly.
7. Conclusions

In this paper, we have discussed different Lagrangian quantities in regimes I, II and III. In regime I, the presented results confirm the general picture of weakly rotating RB convection flow: the flow dynamics are governed by the LSC. The changes in the velocity and acceleration statistics (rms and autocorrelations) both at the cell centre and near the top plate are relatively small in this regime.

According to the heat transfer efficiency, the transition from regime I to regime II is known to be sudden: The LSC is replaced by vortical plumes, in particular close to the wall and in the boundary layers. The rms velocity and acceleration and Lagrangian autocorrelations near the top plate experience a sudden change in trend while crossing this transition. On the other hand, the data at the centre show a gradual change between regime I and II; indicating that the transition starts near the plates, as has been shown earlier in Rajaei et al. (2016a).

The formation of the vortical plumes in regime II and columns in regime III is reflected in the Lagrangian velocity and acceleration statistics. The vertical velocity fluctuations decay faster than their horizontal counterparts at \( z = 0.8H \) and \( z = 0.975H \). The integral time scales show an increase as \( Ro \) decreases at the cell centre and near the top plate, except for the horizontal component near the top plate. The horizontal acceleration fluctuation near the top increases while the other components decrease with decreasing \( Ro \). All these observations are due to the formation of the vortical plumes and columns close to the top plate.

The transition between regimes II and III is comparatively gradual: the Lagrangian velocity statistics show gradual changes. However, the acceleration fluctuations show rather abrupt changes while crossing to regime III, keeping in mind that there is a contribution from inertial waves in the acceleration statistics. The crossover to regime III coincides with the following phenomena and observations:

- **The vertical motions are suppressed**: The vertical component of the rms velocity and acceleration values are suppressed at a higher rate in regime III compared to regime II with decreasing \( Ro \). The suppression of the vertical motions result in a decrease in heat transfer efficiency.

- **The vortical plumes penetrate further into the bulk**: The horizontal component of the rms acceleration continue to increases with decreasing \( Ro \) at \( z = 0.975H \), except for the lowest considered \( Ro \). At \( z = 0.8H \), the horizontal component of the rms acceleration start to increase with decreasing \( Ro \) while transitioning to regime III: indicating that plumes have reached the interior region in regime III. At the cell centre, no enhancement in the horizontal component of the rms acceleration is observed. One should note that acceleration and velocity might not be perfect criteria for evaluating whether plumes reach the cell centre because it is known that as we go down inside a plume the sign of vorticity changes, e.g. a cold plume has a cyclonic vorticity near the top while it possesses an anticyclonic vorticity near the bottom plate. Therefore, the vorticity field is weaker in the cell centre resulting in a weaker swirling motion in the horizontal plane. This vertical change of vorticity can be explained with the thermal-wind balance. As a result, if the horizontal acceleration and velocity statistics at the centre do not follow their counterparts near the top, it does not directly imply that the plumes have not reached the cell centre.

- **The vortical plumes interact less with their surroundings**: Two correlation time scales are present for plume and non-plume regions in regime III which indicates that the flow dynamics has distinct character inside and outside the plumes: the fluid exchange between plume and non-plume regions decreases in regime III for current \( Pr \). In addition, a sudden
increase in the ratio $RA = a_{xy}^{rms}/a_z^{rms}$ at $z = 0.975H$, observed in regime III, supports the same conclusion as the particles inside a plume remain inside the plume and experience a strong horizontal centripetal acceleration.

The first and second observations support the description of the transition to regime III by Julien et al. (2012a). The last observation has not been reported earlier: it might be a key parameter to the transition. However, further experimental and numerical studies are certainly required to evaluate the different mechanisms.

This work is supported by the Nederland Wetenschappelijk Onderzoek (NWO), the Netherlands-formerly known as Stichting voor Fundamenteel Onderzoek der Materie (FOM). R.P.J.K. has received funding from the European Research Council (ERC) under the European Unions Horizon 2020 research and innovation programme (Grant Agreement No. 678634). The authors would like to express special thanks of gratitude to Ad Holten, Gerald Oerlemans, and Freek van Uittert for assisting with the preparation of the experimental setup.

REFERENCES


Velocity and acceleration statistics in rapidly rotating Rayleigh–Bénard convection


