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The influence of the EUV spectrum on plasma induced by EUV radiation in argon and hydrogen gas

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Abstract
Plasmas induced by EUV radiation are scarcely investigated, although they are unique since they are created without any discharge. These plasmas are also of interest from an applicational point of view, since they are related to the lifetime of optics in EUV lithography tools. In order to assess this impact, it is essential to characterize and understand EUV-induced plasma. In this contribution the influence of the background gas (argon and hydrogen) in the lithography tool and the spectrum of the illumination source on the electron density of EUV-induced plasma is investigated using microwave cavity resonance spectroscopy.

The experimental results showed that out-of-band radiation (>20 nm) is the main contributor to EUV-induced plasma in both argon and hydrogen. In hydrogen, this contribution is relatively more important than in argon due to the stronger wavelength dependence of the photoionization cross section of hydrogen than of argon. Furthermore, the production of electrons by out-of-band radiation lasts longer than the production by in-band radiation (10–20 nm) due to the longer temporal width of out-of-band radiation. Finally, the obtained results correspond reasonably well with estimates from a simplified absorption model.

Keywords: EUV, microwave cavity resonance spectroscopy, electron density, spectral influence

(Some figures may appear in colour only in the online journal)
From this the electron density can be determined as \[ n_e = \frac{2m_e\varepsilon_0\omega^2}{e^2}\Delta\omega \] (1)

with \( m_e \) the electron mass, \( e \) the elementary charge and \( \omega_0 \) and \( \omega \) the resonant frequencies without and with plasma, respectively. Note that this method gives the average electron density weighted with the square of the local electric field of the excited mode \( E(\vec{r}) \) [18]:

\[ \overline{n_e} = \frac{\int_{\text{cavity}} n_e(\vec{r})E^2(\vec{r})d\vec{r}}{\int_{\text{cavity}} E^2(\vec{r})d\vec{r}} \] (2)

To determine the local electron density, the spatial profile on the electron density needs to be known. Since this profile is unknown, here we present the square-electric-field-weighted average electron density (\( \overline{n_e} \)).

3. Experimental set-up

Two types of EUV sources have used in this work, a xenon-based and tin-based (laser assisted) discharge produced plasma source. The details of the overall set-up, both sources and the diagnostics are described in this section.

3.1 Overall set-up

Although different EUV sources have been used in this work, the general layout of the experimental set-up is very similar for all experiments (figure 2). The set-up consists of three chambers: the source, collector and measurement chamber. Details of these sources can be found in [21, 22]; the spectra of both sources are given in figure 3. The collector, which is a set of rotationally symmetric multi-shell grazing incidence mirrors located in the collector chamber, collects and focuses the radiation generated by the EUV source in the intermediate focus (IF) in the measurement chamber. The collector and measurement chamber are differentially pumped to allow high vacuum (<10^{-4}\ Pa) in the measurement chamber. An SPF (spectrum purity filter) or aluminium filter can be installed in the beam path; the transmission curves of these filters is given in figure 4. The pressure in the measurement chamber is controlled by a needle valve and stabilized with a feedback loop including the pump system. In all measurements presented here, the pressure was 5 Pa.

3.2 Measurement vessel

The main diagnostic placed in the measurement vessel was microwave cavity resonance spectroscopy (MCRS). In case of the xenon-based source, the cavity is placed around IF; while in case of the tin-based source, the cavity is placed about 20 cm after IF for technical reasons; to ensure that the cavity walls were not illuminated by EUV radiation, an aperture was installed in front of the cavity in tin-based source only. The cavity is made of aluminium and has an inner radius of 33 mm.

![Figure 1. Total photoionization cross sections for both argon [9] and hydrogen [10–12].](image-url)
and an inner height of 20 mm. In order for the EUV beam to go through the cavity, a hole with a diameter of 13 mm is drilled in the top and bottom plate of the cavity. Two copper antennas are placed at opposite sides in the side-wall of the cavity; these antennas are not subject to EUV irradiation. One of these antennas is connected to a low power (10 mW) micro-wave generator (Stanford Research Systems SG386), exciting the TM_{010}-mode around 3.5 GHz in the resonant cavity. The electric field of this mode is calculated using the plasimo platform [26, 27] and is shown in figure 5. This electric field has its maximum value at the axis of the cavity, meaning that the electron density is most effectively sampled in the centre of the cavity (see equation (2)), which is also the radial position at which the EUV beam is directed through the cavity. The other antenna is connected to a microwave detector (Hittite 602LP4E), which has a time response of 10 ns, while the response time of the cavity is 14 ns. The resonant frequency is determined by sweeping the frequency of the microwave generator over a predefined range and measuring the response of the cavity as a function of time. At every point in time, the response of the cavity as a function of frequency is fitted with a Fourier series (two components). The resonant frequency is the frequency where the fit has its maximum. In [5] it has been shown that the response time of the entire detection system is about 17 ns. The error in the square-electric-field-weighted average electron density is less than 30% and the detection limit is $2 \times 10^{12}$ m$^{-3}$ [5].
Figure 5. Modelled electric field of the TM_{010} mode in the cavity. The electric field is calculated with the plasimo platform [26, 27]. Moreover, the electric field of an ideal cavity without holes is shown. This graph is reproduced from [6], copyright 2015 IOP Publishing.

Furthermore, the EUV power is measured with either a temperature sensor in the xenon-base source or the FFST (far field screening tool) in the tin-based source. The power measurement based on the temperature sensor is extensively described in [6]; in this case the power is determined separately for all used filters. The FFST is developed by Xtreme Technologies GmbH to measure the EUV power. At the entrance of the FFST a zirconium filter is installed, which only transmits between 10 and 20 nm. After this filter, a YAG crystal convert the EUV radiation to visible radiation, which is captured by a CCD camera. The intensity of the images recorded by the CCD camera is proportional to the EUV power. To obtain the absolute EUV power, the FFST has been calibrated at a synchrotron facility. The FFST measures the in-band EUV energy; to obtain the total EUV energy (in-band and out-of-band), the in-band EUV energy is extrapolated using the tin spectrum in figure 3. The transmission curves of the filters (figure 4) is used to calculate the effect of the filters on the EUV energy.

4. Results and discussion

The influence of the EUV spectrum on the maximum square-electric-field-weighted average electron density of the EUV-induced plasma in both argon and hydrogen gas was studied using various transmission filters in both a xenon-based and tin-based source.

4.1. Argon gas

4.1.1. Xenon-based source. The measured maximum $n_e$ generated by the xenon-based source with different filters is given in table 1. If a filter is placed in the beam path, the induced maximum $n_e$ is lower since part of the spectrum is blocked by the filter. Although the spectrally integrated EUV pulse energy with the aluminium filter is a factor 11.0 ± 0.8 lower than the EUV pulse energy without filter, the maximum $n_e$ is only a factor 4 ± 2 lower. This is because the photoionization cross section in the 20–70 nm band is much higher than in the 10–20 nm band. The importance of the 20–70 nm band also explains why $n_e$ decreases by a factor 7 ± 3 due to the SPF (which blocks the 20–70 nm band), while the spectrally integrated EUV pulse energy only decreases by a factor 2.0 ± 0.1.

The influence of the filters on photoionization of the gas in the cavity can also be estimated with the applied EUV spectrum and the wavelength dependent photoabsorption cross section, which is extensively described in [6]. As discussed in this reference, the maximum $n_e$ due to photoionization $n_{e,\text{max},\text{pi}}$ is given by:

$$n_{e,\text{max},\text{pi}} = p \int_0^\infty \frac{1.2 \sigma_{\text{pi}}(\lambda) E_{\text{EUV}}(\lambda)}{k_B T_g A_{\text{EUV}} h c \lambda} d\lambda$$

with $p$ the pressure, $\sigma_{\text{pi}}$ is the photoionization cross section (see figure 1), $E_{\text{EUV}}(\lambda)$ is the wavelength $\lambda$ dependent EUV pulse energy, $k_B$ is Boltzmann’s constant, $T_g$ is the gas temperature, $A_{\text{EUV}}$ is the EUV beam cross section (13 mm²), $h$ is Planck’s constant and $c$ is the speed of light. The wavelength dependent EUV pulse energy is obtained by calculating the relative spectrum (figure 3) with a wavelength integrated energy measurement, i.e. the relative spectrum is integrated over the wavelength and divided by the integrated EUV energy, which yields a calibration factor to convert the relative spectral intensity to an absolute energy scale. Using equation (2), the estimated electron density, which is valid in the centre of the cavity, is converted to an estimated $n_e$, assuming the spatial distribution of the electron density is a square with a width equal to the FWHM of the EUV beam (2 mm). The results are also listed in table 1 and show very good correspondence with the measurements. This means that the calculations can be used to estimate the contributions of various bands of EUV radiation to the electron density (table 2).

Table 1. Measured and estimated maximum $n_e$ in 5 Pa argon plasma generated by a xenon-based EUV source.

<table>
<thead>
<tr>
<th>Filter</th>
<th>Energy factor $n_e$ (m⁻³)</th>
<th>Measurement $n_e$ (m⁻³)</th>
<th>Estimation</th>
</tr>
</thead>
<tbody>
<tr>
<td>None</td>
<td>(8 ± 2) x 10¹⁵</td>
<td>1 x 10¹⁶</td>
<td></td>
</tr>
<tr>
<td>SPF</td>
<td>2.0 ± 0.1, (1.1 ± 0.3) x 10⁷ ± 3</td>
<td>9 x 10⁷ 11</td>
<td></td>
</tr>
<tr>
<td>Aluminium</td>
<td>11.0 ± 0.8, (1.9 ± 0.6) x 10⁴ ± 2</td>
<td>2 x 10⁴ 5</td>
<td></td>
</tr>
</tbody>
</table>

4.1.2. Tin-based source. The measured maximum $n_e$ generated by a tin-based source with different filters is given in table 3. By applying an SPF the density decreases by a factor 3 ± 1, while the EUV pulse energy decreases by a factor 2.3. The aluminium filter decreases the EUV pulse energy with
For the tin-based source \( n_e \) can also be estimated from the spectrum and absorption cross section as described in the previous paragraph. In this case the spectrum of the tin source was used (figure 3). In order to obtain the spectrum of the radiation entering the cavity, the spectrum of the light generated by the source is multiplied by the transmission of the collector and the transmission of the argon gas between the source and the cavity (about 3 m). The resulting spectrum is calibrated with a wavelength integrated energy measurement of the EUV source. If a filter is placed in the beam path, the spectrum is also multiplied with its transmission. The EUV beam has a radius of 6.4 mm in the centre of the cavity. Using equation (2), the estimated maximum electron density is converted to the maximum \( n_e \). To this end, it is assumed that the spatial distribution of the electron density is a square with a width equal to the diameter of the EUV beam. The estimated \( n_e \) is about a factor 3 larger than the measured \( n_e \). However, the estimated \( n_e \) difference due to the filters corresponds to the measured factors. Hence, the difference between the estimated and measured \( n_e \) is probably caused by an error in the EUV pulse energy estimation or due to the conversion from maximum electron densities to maximum \( n_e \). This means that also for this source, the calculations can be used to estimate the relative contributions of various bands of EUV radiation (table 4). The 20–70 nm band radiation only contributes to 10% of the EUV-induced plasma, while the 10–20 nm band generates 90% of the plasma. In the previous section it was shown that the radiation above 50 nm is the most important source for photoionization. Since in the tin-based EUV source radiation above 30 nm is very small, also the photoionization

<table>
<thead>
<tr>
<th>Filter</th>
<th>Energy factor</th>
<th>Measurement ( n_e (m^{-3}) )</th>
<th>Estimation ( n_e (m^{-3}) )</th>
</tr>
</thead>
<tbody>
<tr>
<td>None</td>
<td></td>
<td>((2.5 \pm 0.8) \times 10^{15})</td>
<td>(8 \times 10^{15})</td>
</tr>
<tr>
<td>SPF</td>
<td>2.3</td>
<td>((8.6 \pm 3) \times 10^{14} 3 \pm 1)</td>
<td>(3 \times 10^{15} )</td>
</tr>
<tr>
<td>Aluminium</td>
<td>7.7</td>
<td>((4.4 \pm 1) \times 10^{14} 6 \pm 2)</td>
<td>(1 \times 10^{15} )</td>
</tr>
</tbody>
</table>

Table 3. Measured and estimated maximum square-electric-field-weighted average electron density in 5 Pa argon generated by a tin-based EUV source.

The influence of the different filters on the measured \( n_e \) can also be estimated theoretically from the filtered EUV spectrum and the photoionization cross section as discussed in [7] and the previous section:

\[
\frac{1.055 \sigma_{\text{p}}(\lambda) E_{\text{EUV}}(\lambda)}{k_B T_e \lambda^2 A_{\text{EUV}} h c / \lambda} d\lambda.
\]  

(4)

The relevant estimates are also listed in table 5 and they correspond very well with the measurements. These calculations can thus be used to estimate the contributions of various EUV bands to \( n_e \) (see table 6). Just as in the case of argon, the 20–70 nm band is most important for the initial generation of plasma (96%) while most energy is within the 10–20 nm band.

4.2. Hydrogen gas

4.2.1. Xenon-based source. The measured maximum \( n_e \) in an EUV-induced hydrogen plasma using a xenon-based source with different filters (SPF and aluminium) are given in table 5. The EUV pulse energy at the entrance of the cavity is a factor 11.0 ± 0.8 lower if an aluminium filter (transmission between 20–70 nm) is placed in the beam path. However, since the photoionization cross section is higher in the 20–70 nm band than in the 10–20 nm band, \( n_e \) is only a factor 4 ± 2 lower. This also explains why \( n_e \) decreases with a factor 27 ± 12 if an SPF (transmission between 10 and 20 nm) is used, while the EUV pulse energy only decreases by a factor 2.0 ± 0.1.

The influence of the different filters on the measured \( n_e \) can also be estimated theoretically from the filtered EUV spectrum and the photoionization cross section as discussed in [7] and the previous section:

\[
\frac{1.055 \sigma_{\text{p}}(\lambda) E_{\text{EUV}}(\lambda)}{k_B T_e \lambda^2 A_{\text{EUV}} h c / \lambda} d\lambda.
\]  

(4)

The relevant estimates are also listed in table 5 and they correspond very well with the measurements. These calculations can thus be used to estimate the contributions of various EUV bands to \( n_e \) (see table 6). Just as in the case of argon, the 20–70 nm band is most important for the initial generation of plasma (96%) while most energy is within the 10–20 nm band.

Furthermore, the temporal evolution of the creation of EUV-induced plasma is affected by the various filters (figure 6). When an SPF is placed in the beam path, the maximum \( n_e \) is reached earlier than without any filter or with an aluminium
filter. This is caused by the difference in temporal evolution of the EUV spectrum. The various bands of the EUV spectrum have a different temporal width. The temporal width of the EUV pulse filtered with an SPF is shorter than without any filter (figure 6), i.e. the 10–20 nm band has a shorter temporal width than the 20–70 nm band. This is why the production of electrons due to photoionization lasts longer without any filter or with an aluminium filter; hence, the maximum $n_e$ is reached later.

### 4.2.2. Tin-based source

The measured maximum $n_e$ in an EUV-induced plasma in hydrogen using a tin-based source with an SPF and aluminium filter are given in table 7 along with the theoretical predictions based on the applied EUV spectrum and photoionization cross section. The effect of the filters on the electron density is captured well by the theoretical estimation. Hence, these calculations can be used to estimate the contribution of various bands of the EUV spectrum to the total $n_e$ (table 8). The 10–30 nm band contributes to 96% of $n_e$. In this band, most of the EUV pulse energy is in the 10–20 nm range. However, the 4% of EUV pulse energy which is in the 20–30 nm range still contributes 41% of $n_e$. This is because the higher photoionization cross section compensates for the lower EUV pulse energy in the 20–30 nm band. This means that by controlling the spectrum of the EUV source, the plasma density can also be controlled.

If the spectral influence in a hydrogen EUV-induced plasma is compared with an argon EUV-induced plasma, some significant differences can be observed. The 20–30 nm-band is more important for generation of electrons in hydrogen (41%) than it is in argon (9%). This is because the photoionization cross section of hydrogen depends more strongly on the wavelength than the photoionization cross section of argon (figure 1). As a result, the 20–30 nm-band has a larger contribution to $n_e$ in hydrogen than in argon. The effect of electron impact ionization is too small to explain this difference [6, 7]. Furthermore, the dependence of the electron impact ionization cross section on the electron energy is alike in argon and hydrogen [28, 29].

### 4.3. Comparison of the Xe- and Sn-based source

The main difference between the xenon- and tin-based EUV source is the generated spectrum. In short, the xenon-based source contains significantly more radiation above 20 nm than the tin-based source. Since the photoionization cross section for this part of the spectrum is higher, the EUV-induced plasma will be more dense even while the total EUV pulse energy is the same. This means that not only the total EUV pulse energy is important for creation of EUV-induced plasma, but even more the spectrum of the source. Since the radiation above 50 nm is most important for photoionization, $n_e$ of the EUV-induced plasma is reduced drastically by reducing out-of-band radiation.

A significant difference in the decay of EUV-induced plasma generated by a xenon- or tin-based EUV source was observed. Although it has not been verified yet, it should be pointed out that the gas flows in the two systems might be different, which could influence the decay.
5. Conclusions

In this contribution, the influence of the EUV spectrum on a pulsed EUV-induced plasma in argon and hydrogen is studied. It was found that out-of-band radiation (>20 nm) is the main contributor to EUV-induced plasma. In-band radiation (10–20 nm) is less significant. In hydrogen, out-of-band contribution is relatively more important than in argon due to stronger wavelength dependence of the photoionization cross section of hydrogen than of argon. Due to the larger temporal width of out-of-band radiation, the production of electrons lasts longer due to out-of-band radiation than due to in-band radiation. Furthermore, a simplified absorption model can be used to obtain reasonable estimates of the electron density.

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References

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